

An artistic rendering of the Huygens probe in orbit around Saturn. The probe is shown in the center, with its gold-colored body and various instruments. A large, white, parabolic dish antenna is extended from the probe. The background features the bright, yellowish-orange surface of Saturn and its prominent rings, set against a dark, star-filled space. The Earth's horizon is visible on the left side of the image, showing a blue and white atmosphere over a brownish-orange surface.

# HUYGENS RADIO LINK IN-FLIGHT PERFORMANCE And some science results

Miguel Pérez Ayúcar, ESTEC/ESA  
IPPW3, Annavyssos, Greece 26 Jun 05

# HUYGENS RADIO LINK IN-FLIGHT PERFORMANCE

And some science results

## OUTLINE of the presentation

- 1.- HUYGENS data relay link description
- 2.- Link design flaw discovery and recovery mission
- 3.- In-flight link reconstruction and performance overview
- 4.- Some science aspects/results derived from the engineering link parameters
- 5.- Conclusions

Back-up slides



# HUYGENS RADIO LINK IN-FLIGHT PERFORMANCE

And some science results

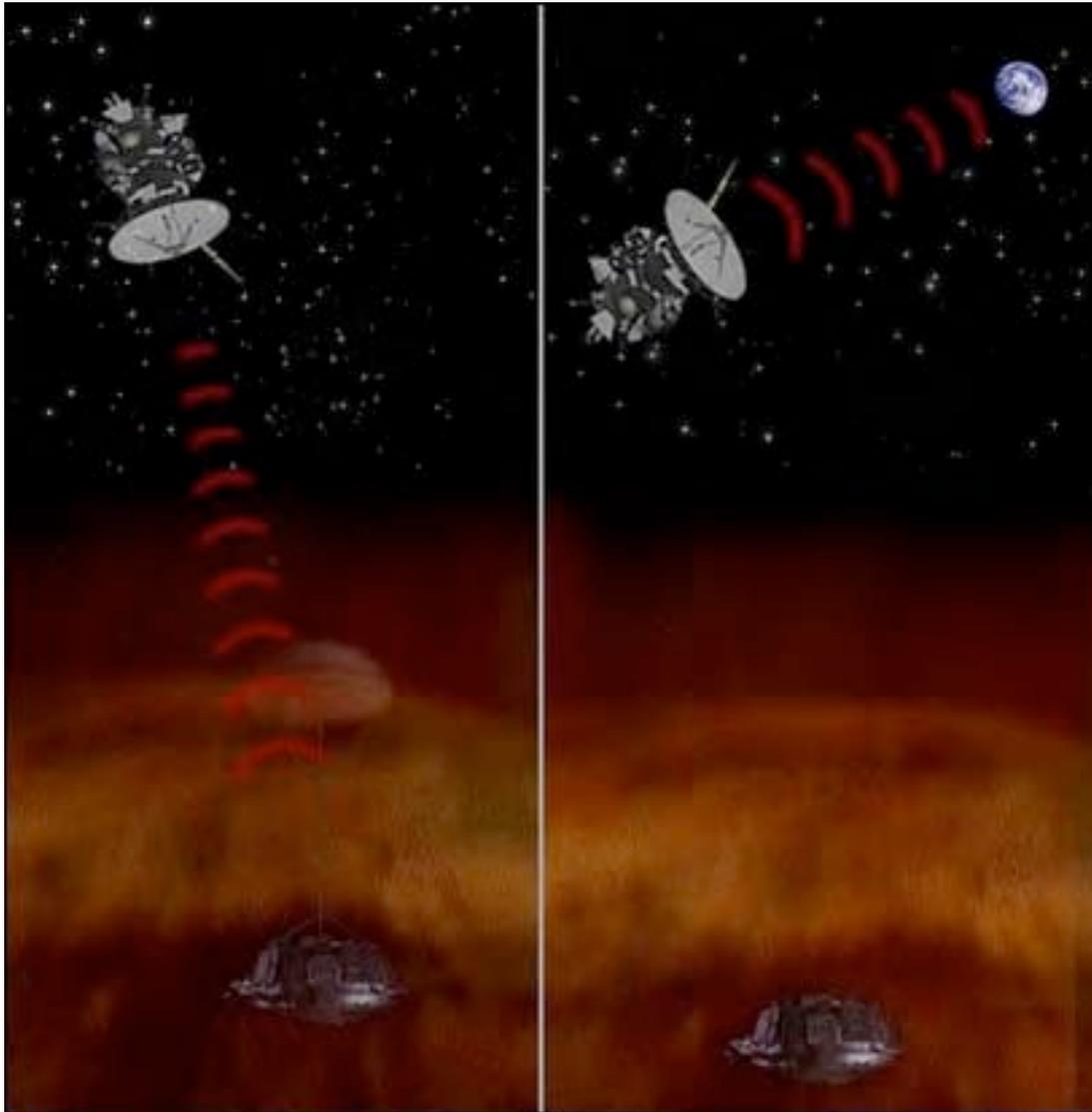
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# THE HUYGENS DATA RELAY LINK



- 2 CHAINS (redundancy)
- PCM/BPSK/PM ONE-WAY:
  - ChA: 2040 MHz LHCP
  - ChB: 2080 MHz RHCP
- Residual carrier phase modulation ( $m=1.34, 1.37$ )
- Subcarrier 131072 Hz, BPSK modulated
- Data PCM-NRZ-M 8kbps, Reed-Solomon + Convolutional CODE (2:1) → 16ksps
- PTA antenna: resonant quadrifilar helix antenna



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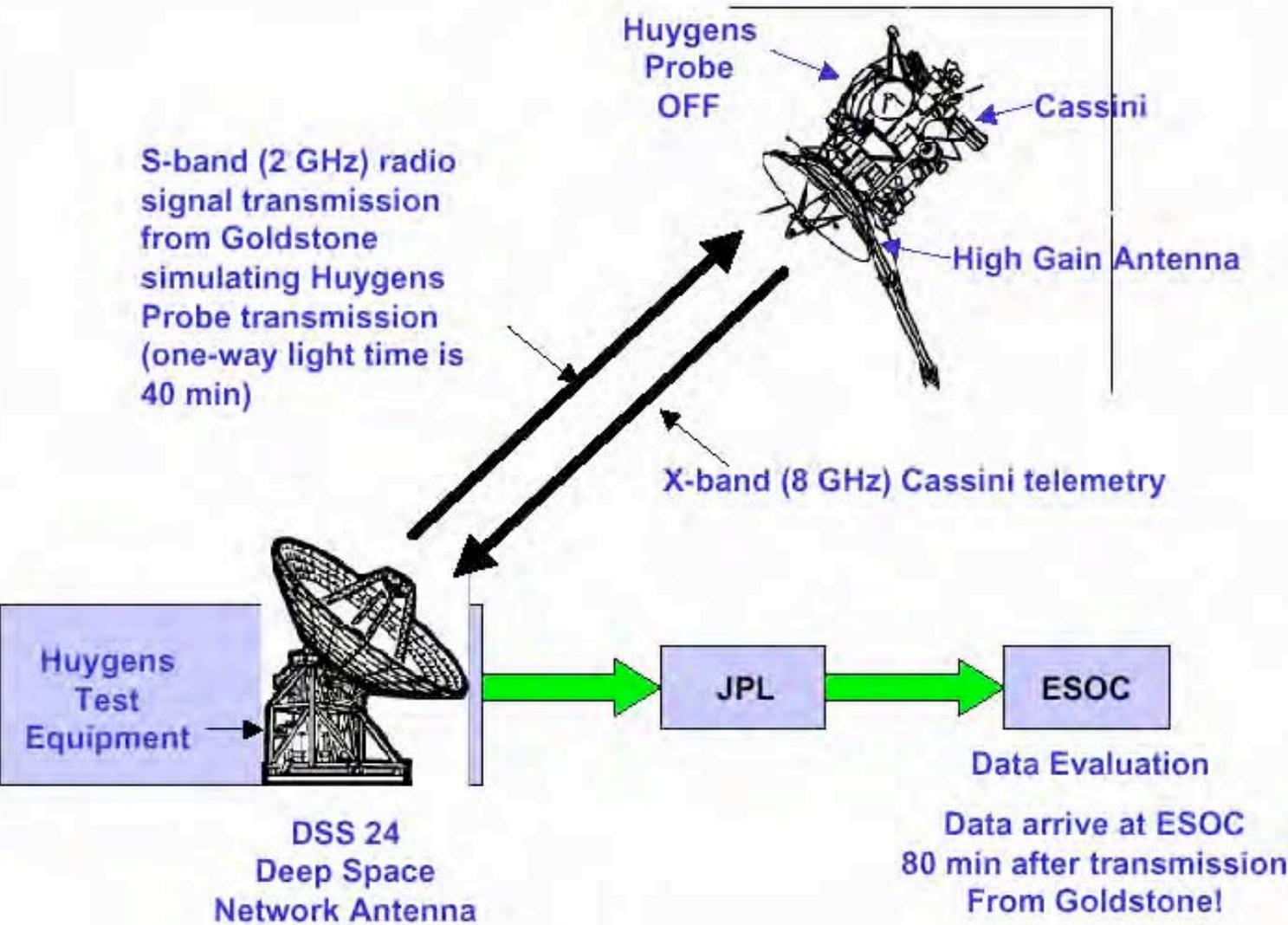
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# End-to-end test: PROBE RELAY TEST (PRT #1) Feb 2000



## **PRT#1 RESULT:**

**For the expected mission frequency  
Doppler shift, 80-90% of data packets  
would be rejected on-board!!! On Earth,  
severe (CRITICAL) science data loss.**



## **HRTF: Huygens Recovery Task Force**

**~1.5 years, Feb 2000- July 2001**

**Goal: understand flaw and find solutions**

## **HIT: Huygens Implementation Team**

**~3 years, July 2001 – mid-2004**

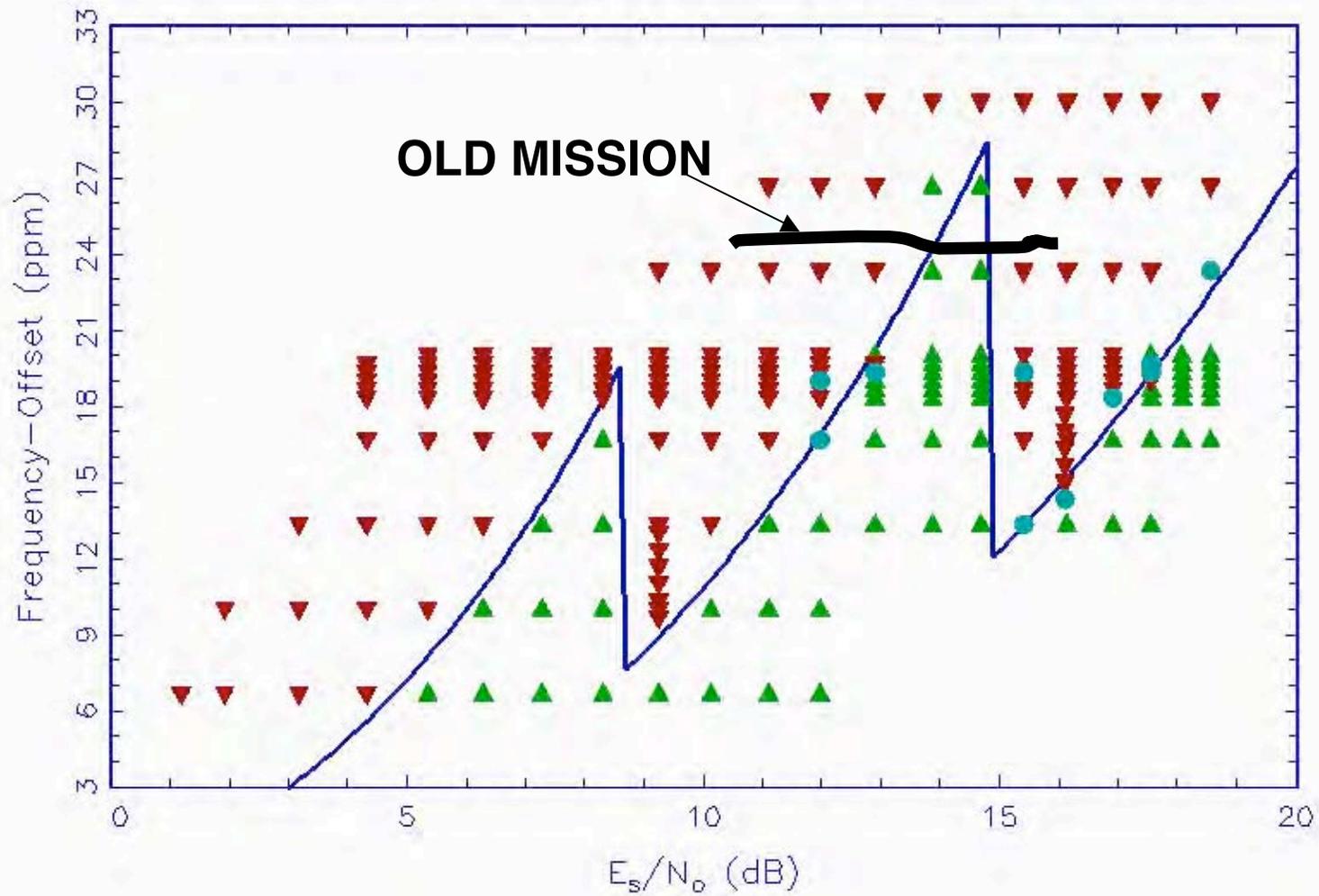
**Goal: implement the recovery solution**



# RECEIVER MODEL and IN-FLIGHT TESTING

GAU55 LPop Thu Jul 12 15:47:15 2001

AF7; USO On; Transition Density: 52.3%; Mod-Index=1.34



## **RECOVERY SCENARIO:**

### **1.- Trajectory change Cassini – Huygens**

**Objective: reduce relative doppler offset**

### **2.- Pre-heating of the probe (4 hours)**

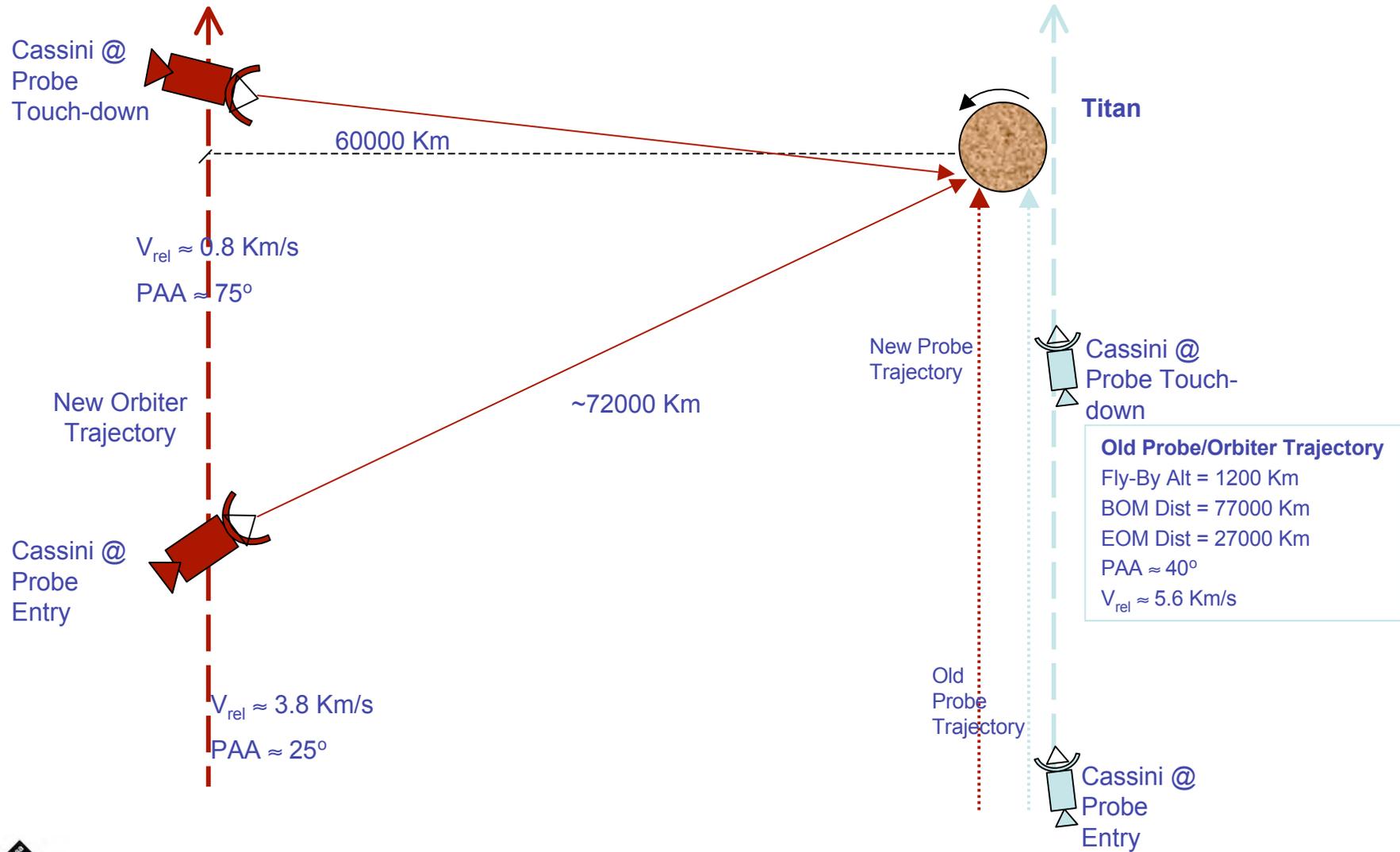
**Objective: decrease the frequency of the data by warming the data stream clock before entry**

### **(3.-) Data post-processing on Earth**

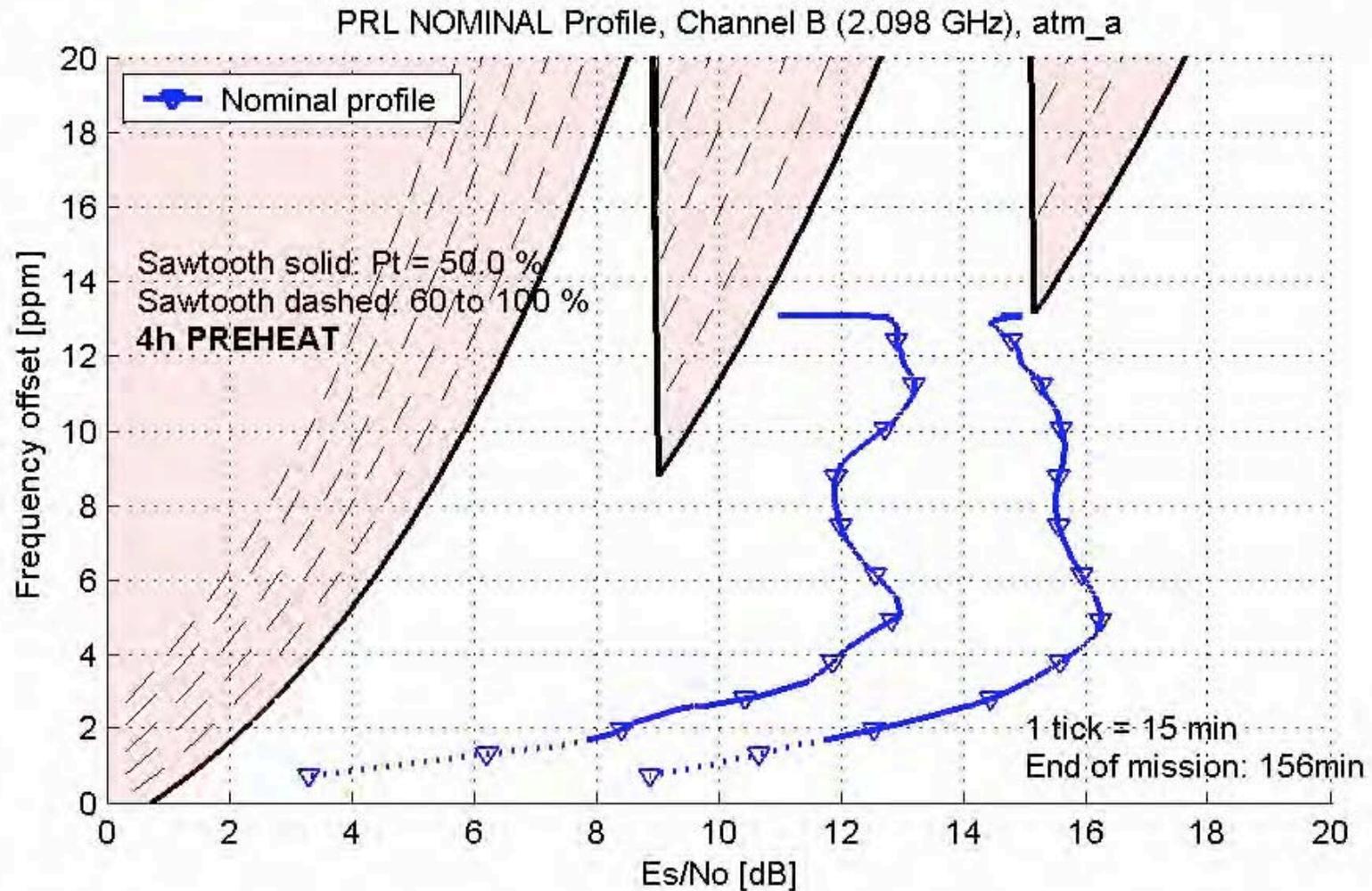
**Further use of Reed-Solomon capability**



# Geometry change



# Predicted performance of the recovery scenario



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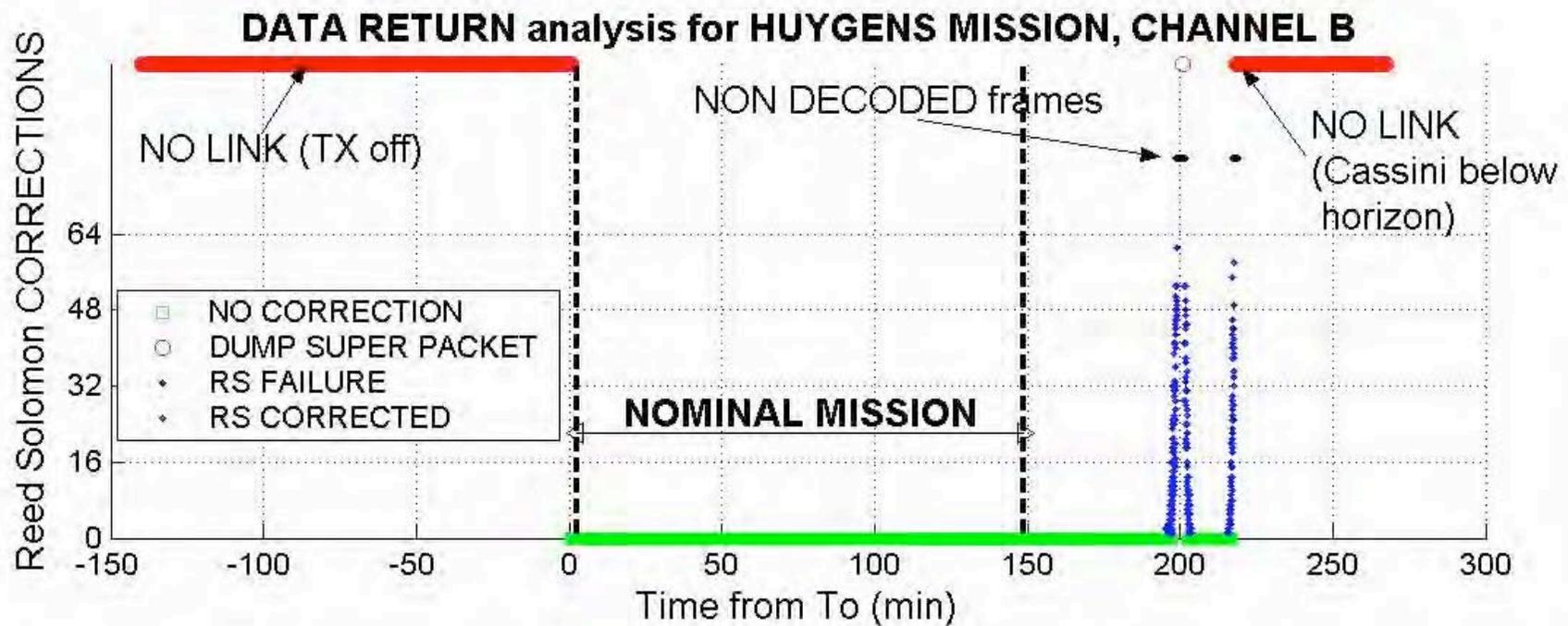
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Back-up slides



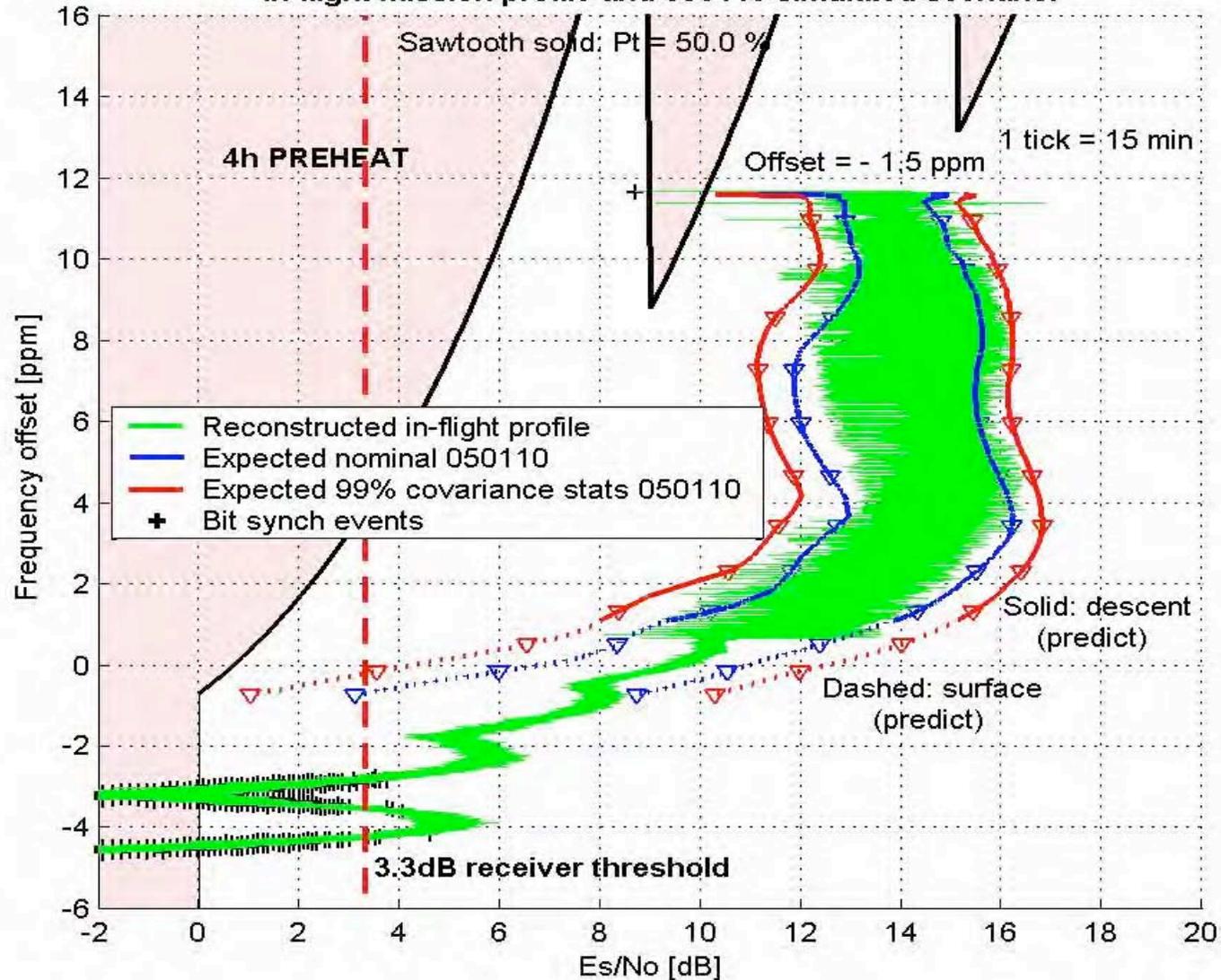
# Data Link PERFORMANCE overview

**100 % DATA RETURN for CH B !! GREAT SUCCESS**  
**+71 min on SURFACE !**



# Finger Plot RECONSTRUCTION

HUYGENS PROBE RELAY RADIOLINK IN-FLIGHT RECONSTRUCTION: Channel B (2.098 GHz).  
In-flight mission profile and 050110 simulated scenario.



# HUYGENS RADIO LINK IN-FLIGHT PERFORMANCE

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## OUTLINE of the presentation

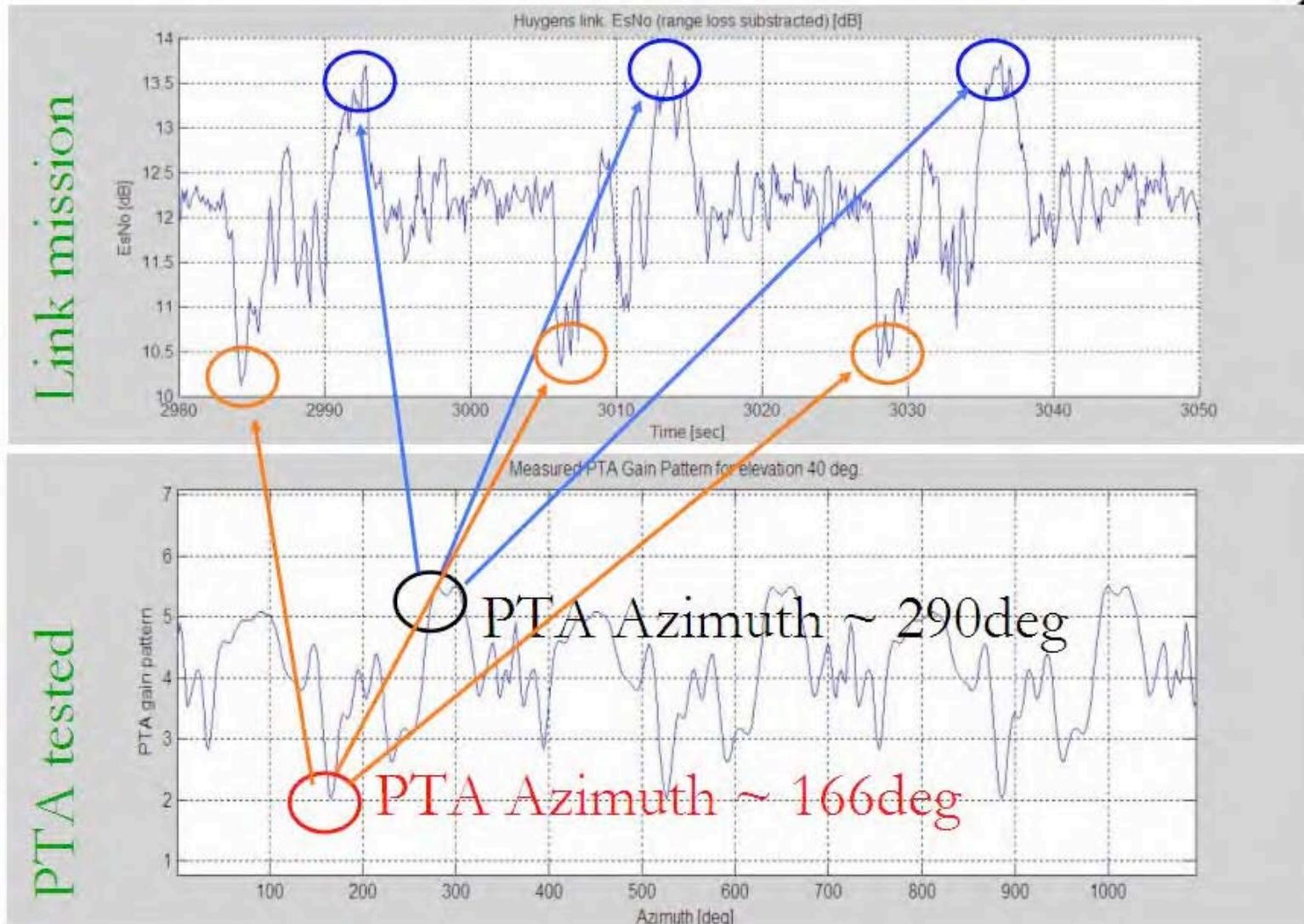
- 1.- HUYGENS data relay link description
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- 4.- Some science aspects/results derived from the engineering link parameters ←
  - (a) Azimuth reconstruction
  - (b) Spin rate reconstruction
  - (c) Link on the surface: radio science
- 5.- Conclusions



Back-up slides

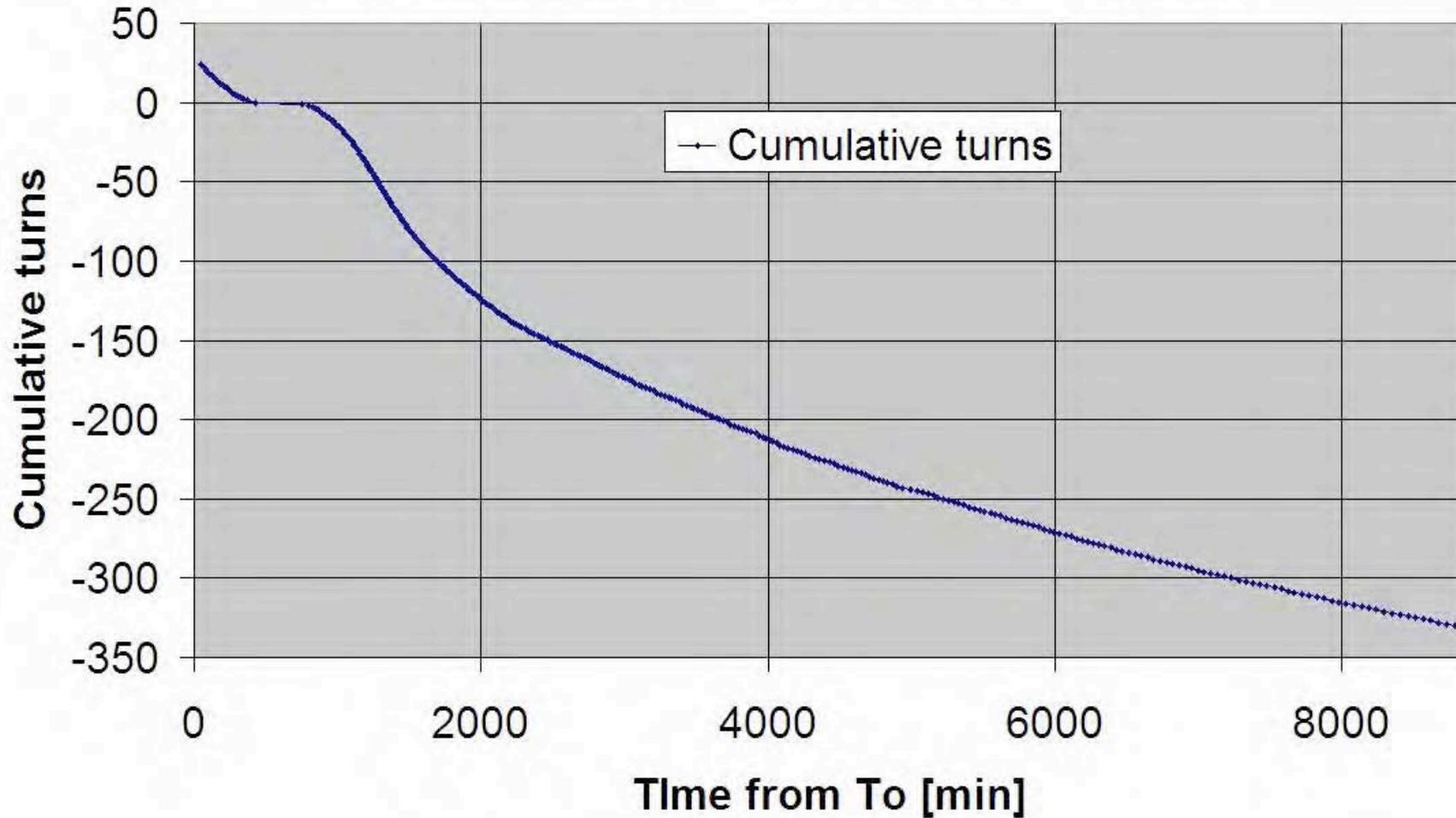
# (a) Absolute AZIMUTH reconstruction:

based on matching the PTA test pattern and the AGC in-flight signal



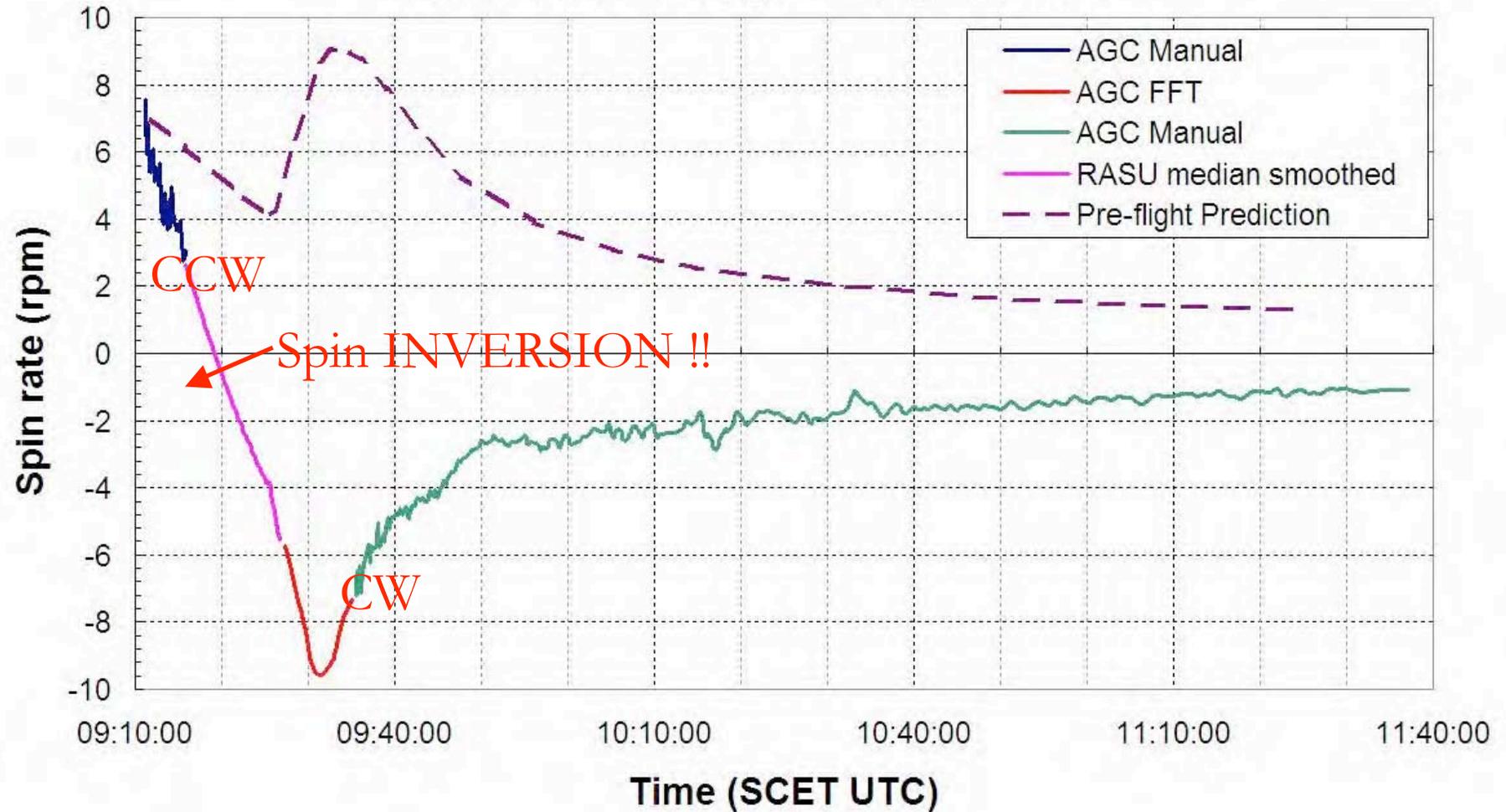
# (a) Absolute AZIMUTH reconstruction

Huygens in-flight data analysis. Cumulative turns

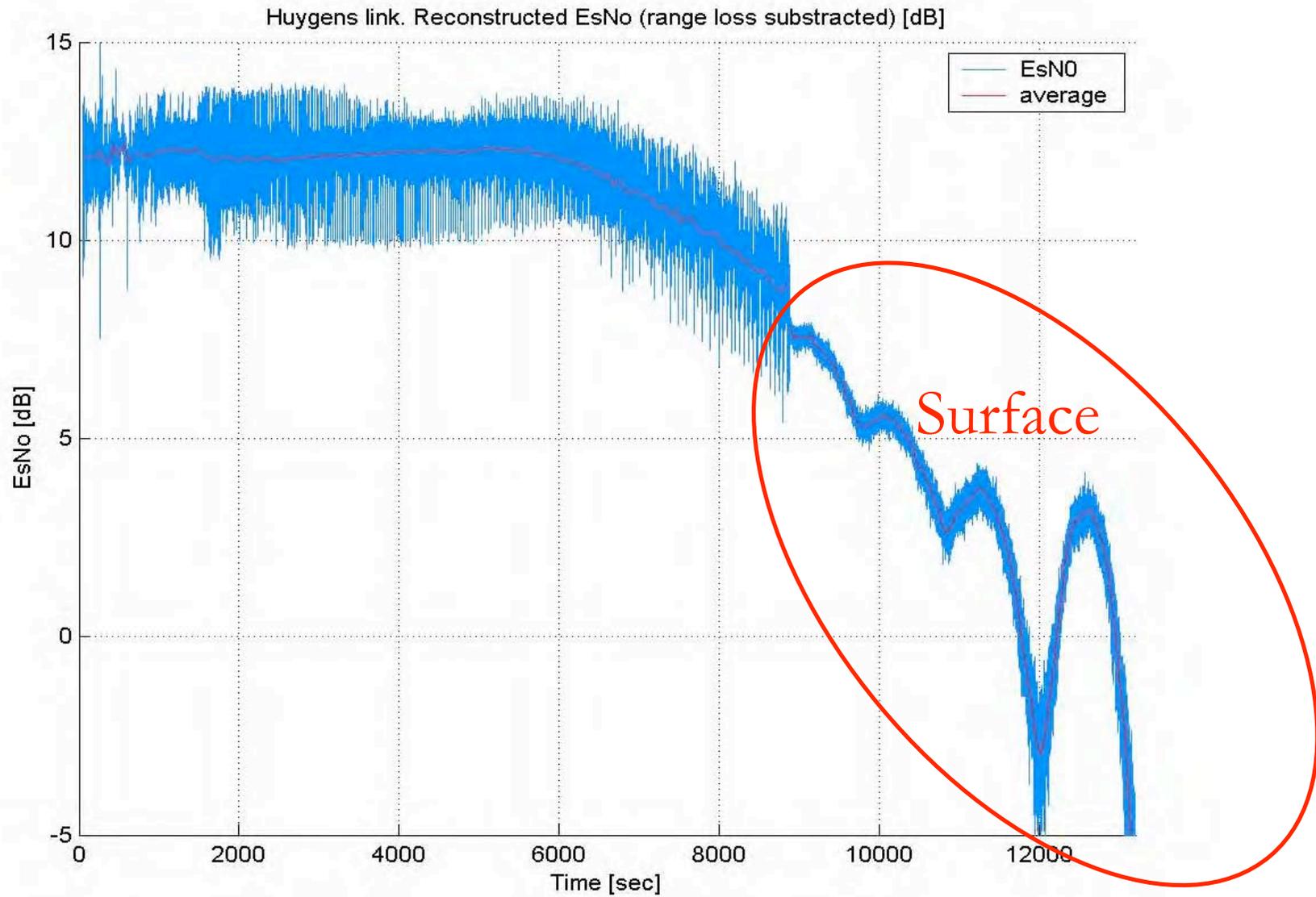


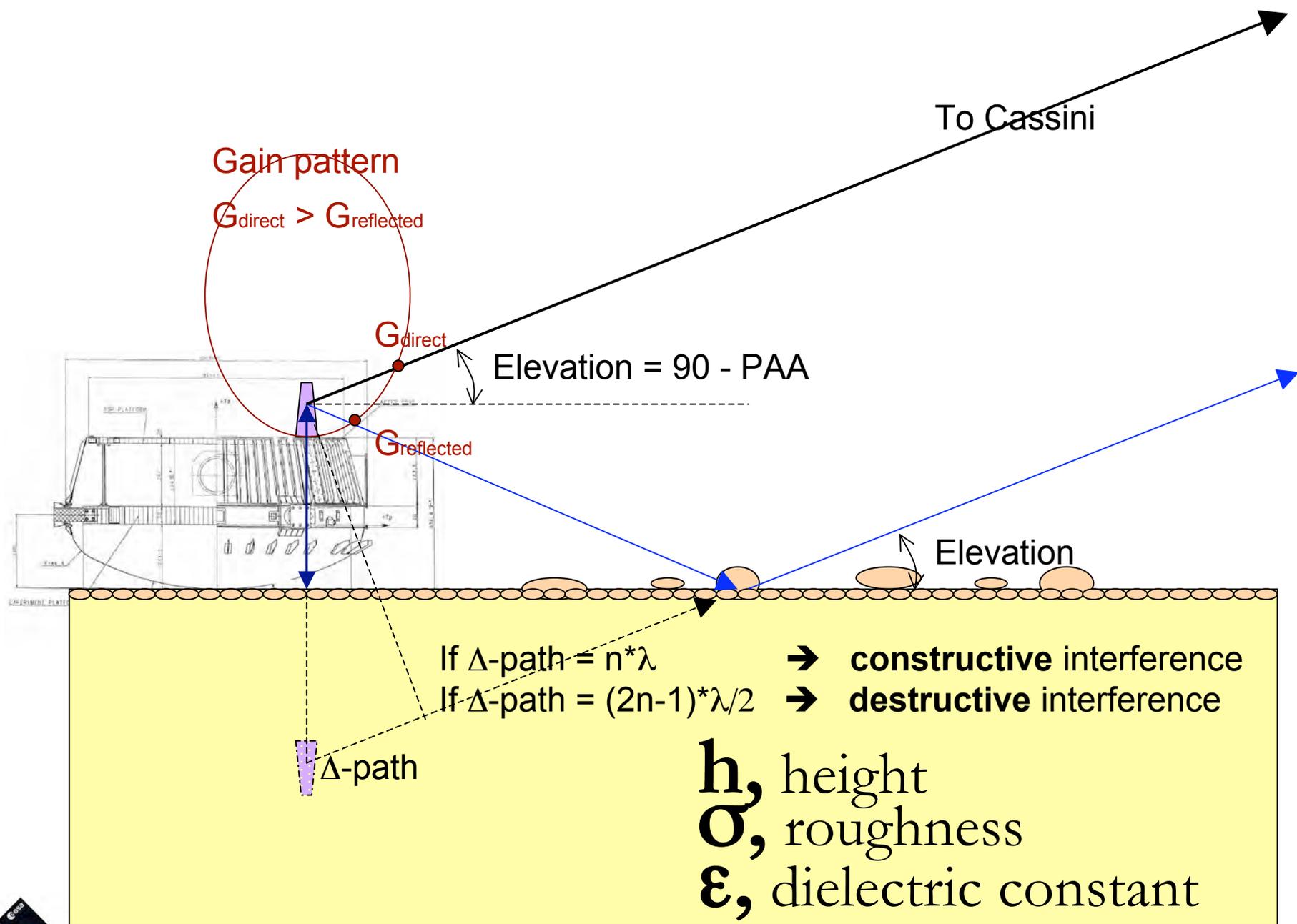
## (b) Spin profile reconstruction

Huygens in-flight spin profile  
derived from engineering sensors

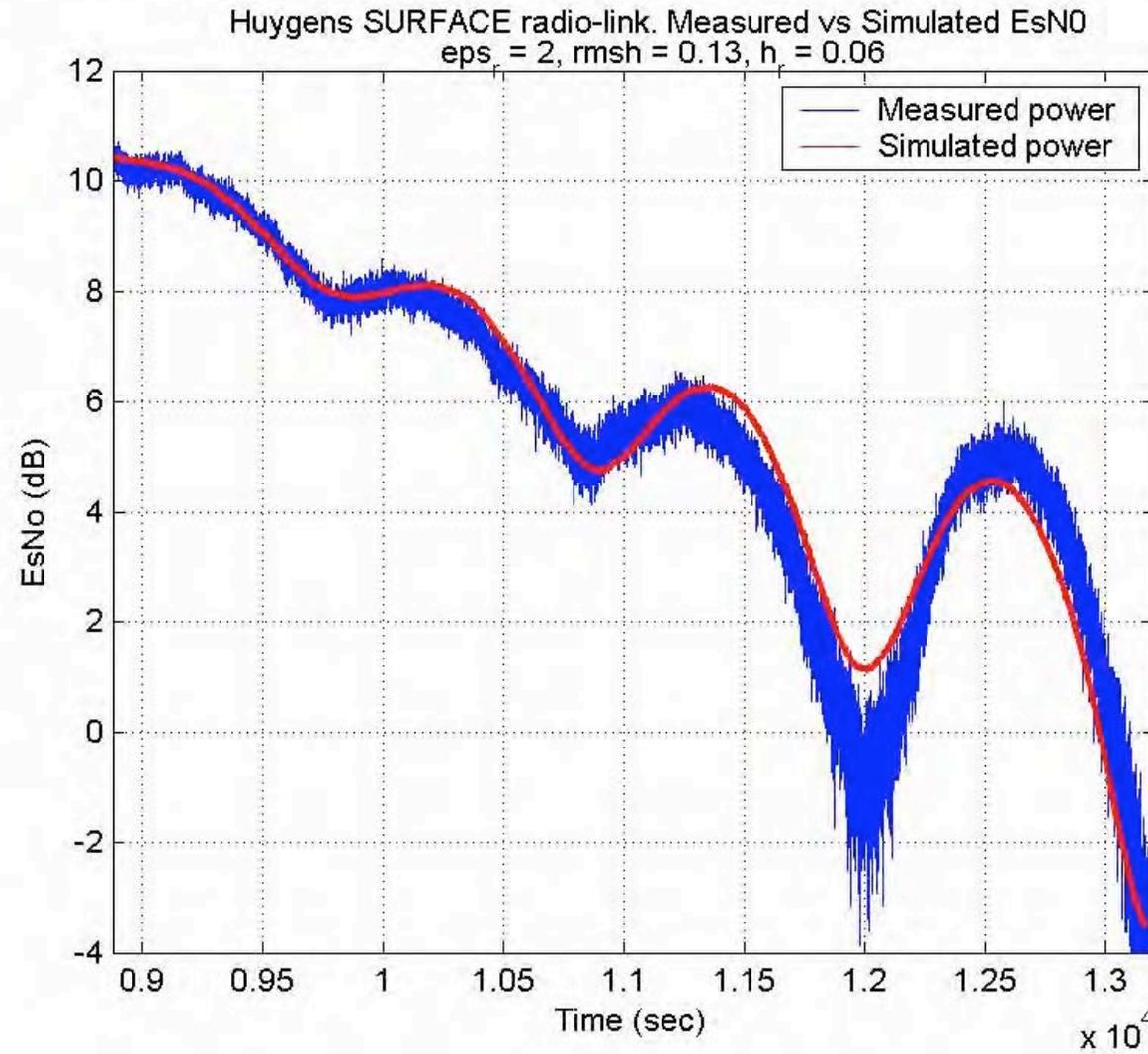


# (c) Link on the Surface





# Simulations



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Back-up slides



# CONCLUSIONS

1.- **Excellent engineering performance** of the Huygens ChB radio link:

- full **100% data returned** back to Earth
- **71 min** 'bonus' on the **surface**, until Cassini set beneath horizon.

Recovery Mission efforts paid off. THANKS HRTF and HIT !!

SCIENTISTS have their data to analyse TITAN.

2.- Engineering telemetries CAN be used for **scientific purposes**, beyond the regular engineering service intended for. BUT need an intelligent planning. Examples:

- The **spin reversal** anomaly has been confirmed and **characterized**.
- **Azimuth and spin rate** profiles have been generated and will help the instrument teams to better interpret their measurements.
- The **multi-path radio-science** behaviour **on the surface** will also

provide a local characterization of the **soil properties** (dielectric constant, roughness) for synergies with the lower res Cassini radar mapping of Titan.

3.- CONGRATULATIONS to all HUYGENS and CASSINI TEAMS for the recovery efforts, and such an awesome mission



# Titan

Distance: 7,298.1 km  
Radius: 2,575.0 km  
Apparent diameter: 30° 14' 10.7"  
Day length: 15.945 days  
Temperature: 89 K

2005 01 14 13:27:02 WEST  
Real time

HUYGENS LANDING SITE

Titan

Phoebe

Betelgeuse

Tethys

Calypso

Pandora

Telesto

Enceladus

Epimetheus

Saturn

Alnilam

# QUESTIONS...

Pan

Dione Rhea

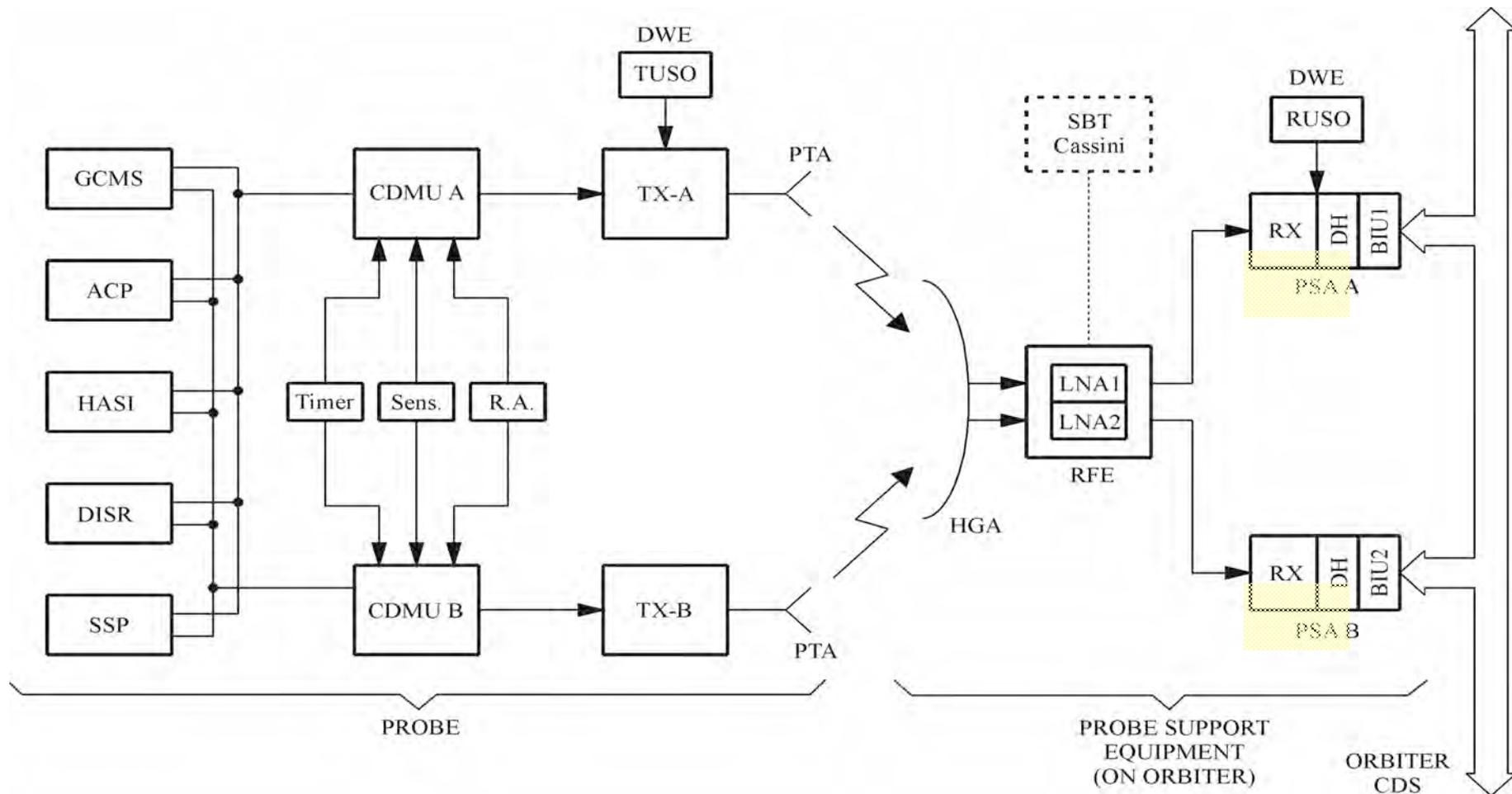
Speed: 0.000 m/s

Sync Orbit Titan  
FOV: 34° 09' 32.0", (1.00x)

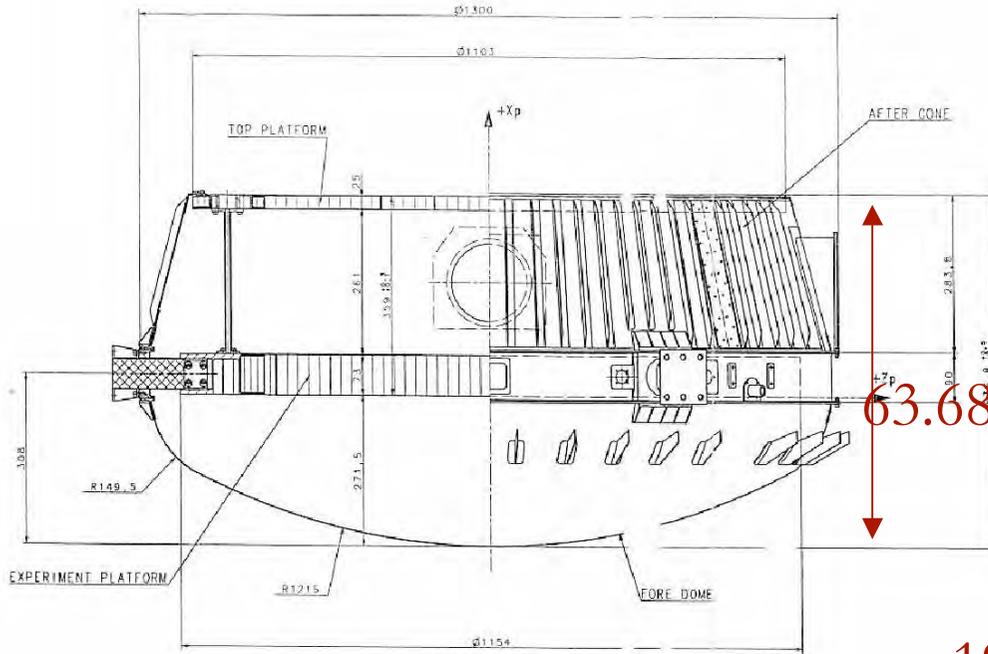
# BACK-UP SLIDES



# Probe Data Relay Subsystem (PDRS)



$$\text{PTA height} = \text{DM height} + \text{PTA phase center} = 63.68 + 19.75/2 = 73.555 \text{ cm}$$

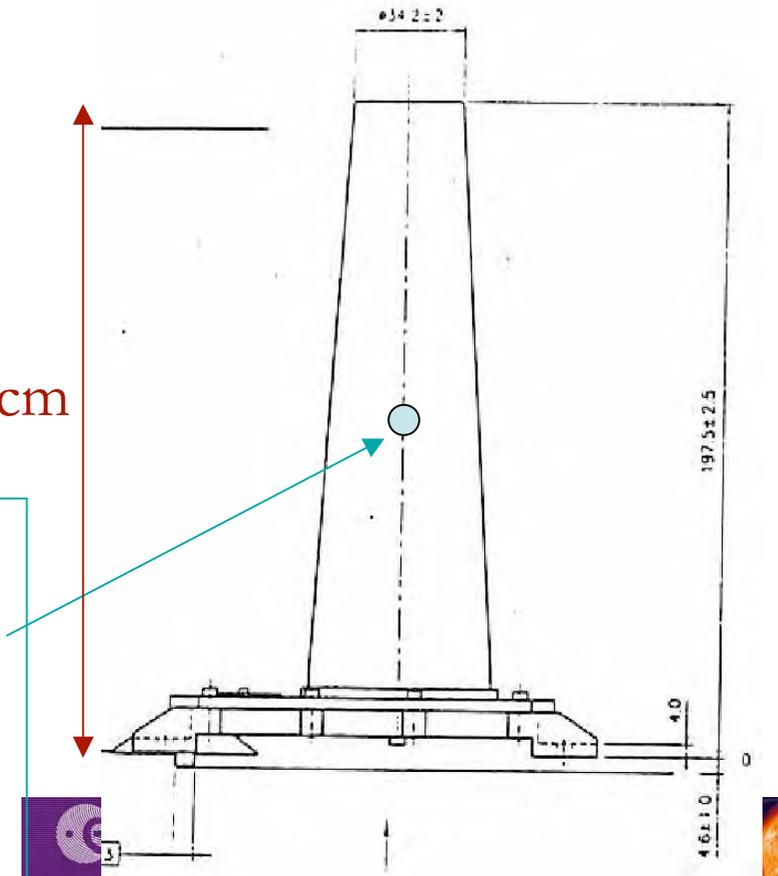


63.68 cm

19.75 cm

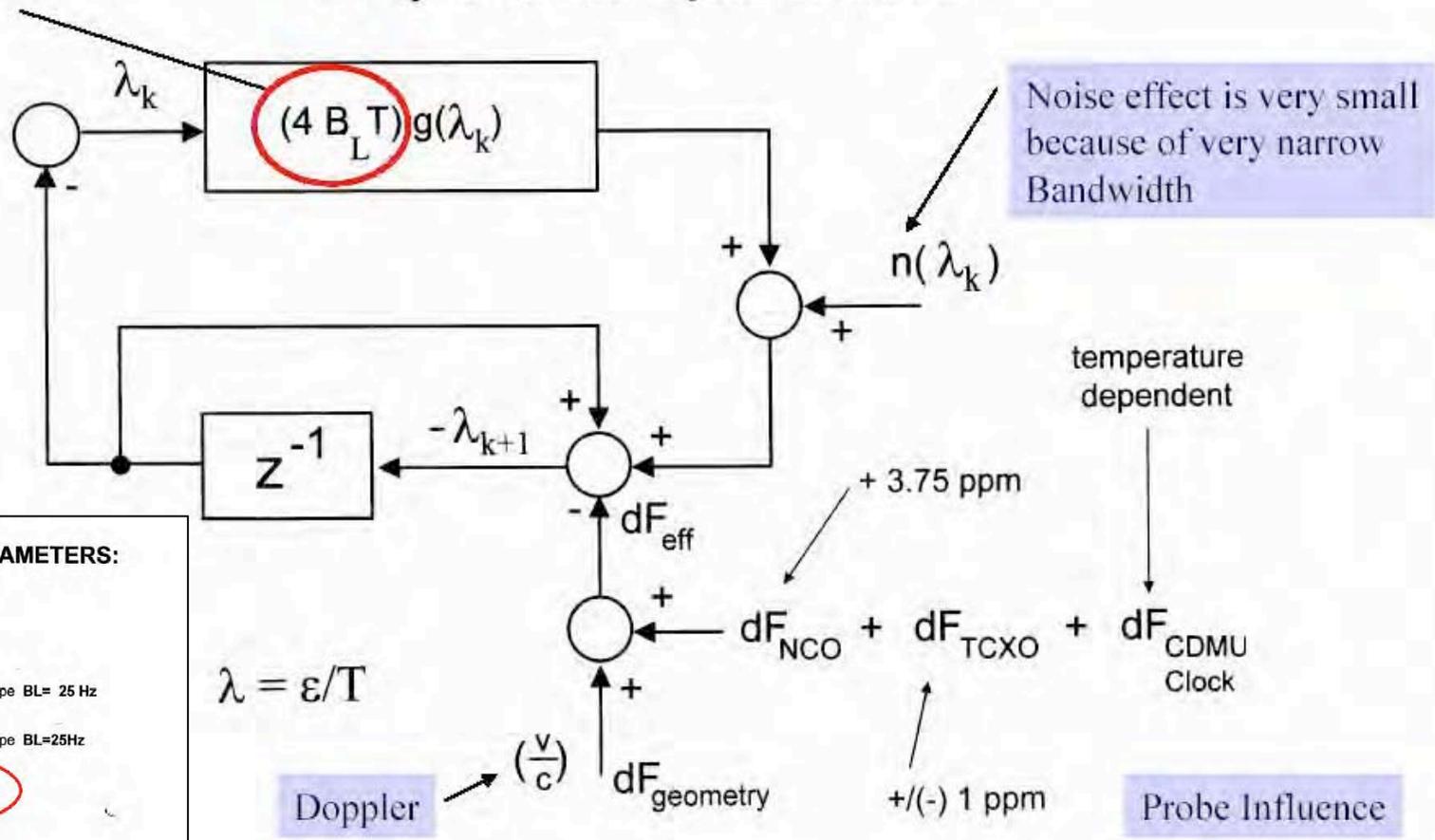
PTA phase center:

Assumed in the middle of the radome for a 4 wires helical antenna (Klooster, ESTEC).



too small

## Bit Synchronizer Dynamic Model



### RECEIVER MAIN PARAMETERS: Tracking Phase

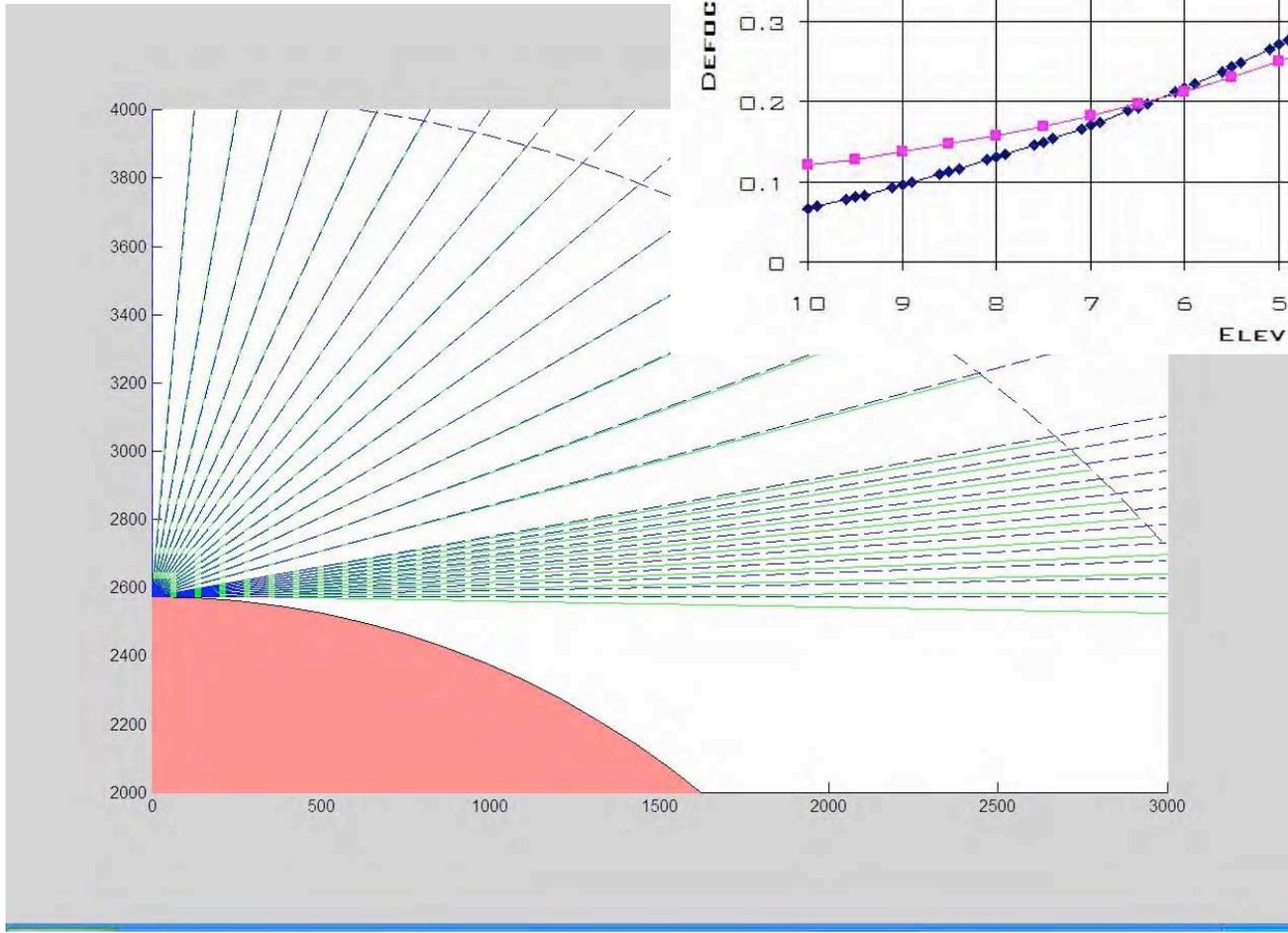
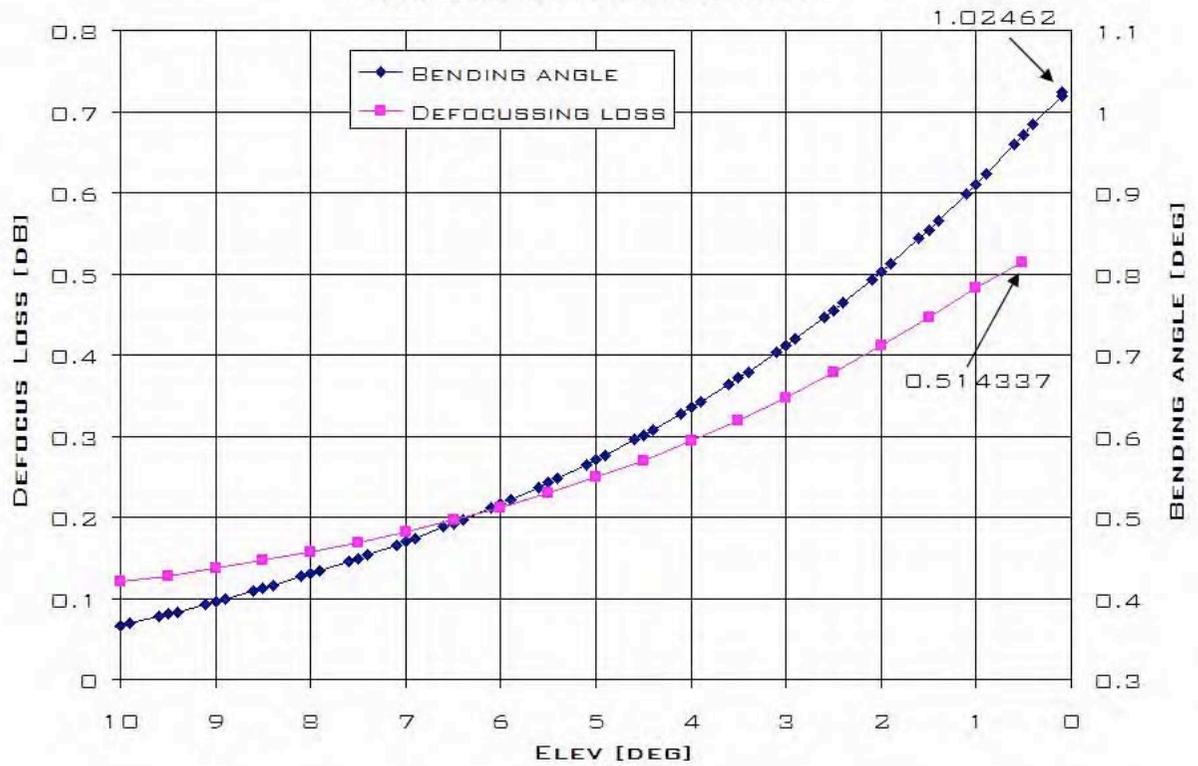
- Carrier Loop
  - second order loop second type BL= 25 Hz
- Subcarrier Loop
  - second order loop second type BL=25Hz
- Bit synch. Loop
  - first order loop BL =3- 6 Hz

Intention, but not satisfied in actual implementation

HUY ROME March, 2000

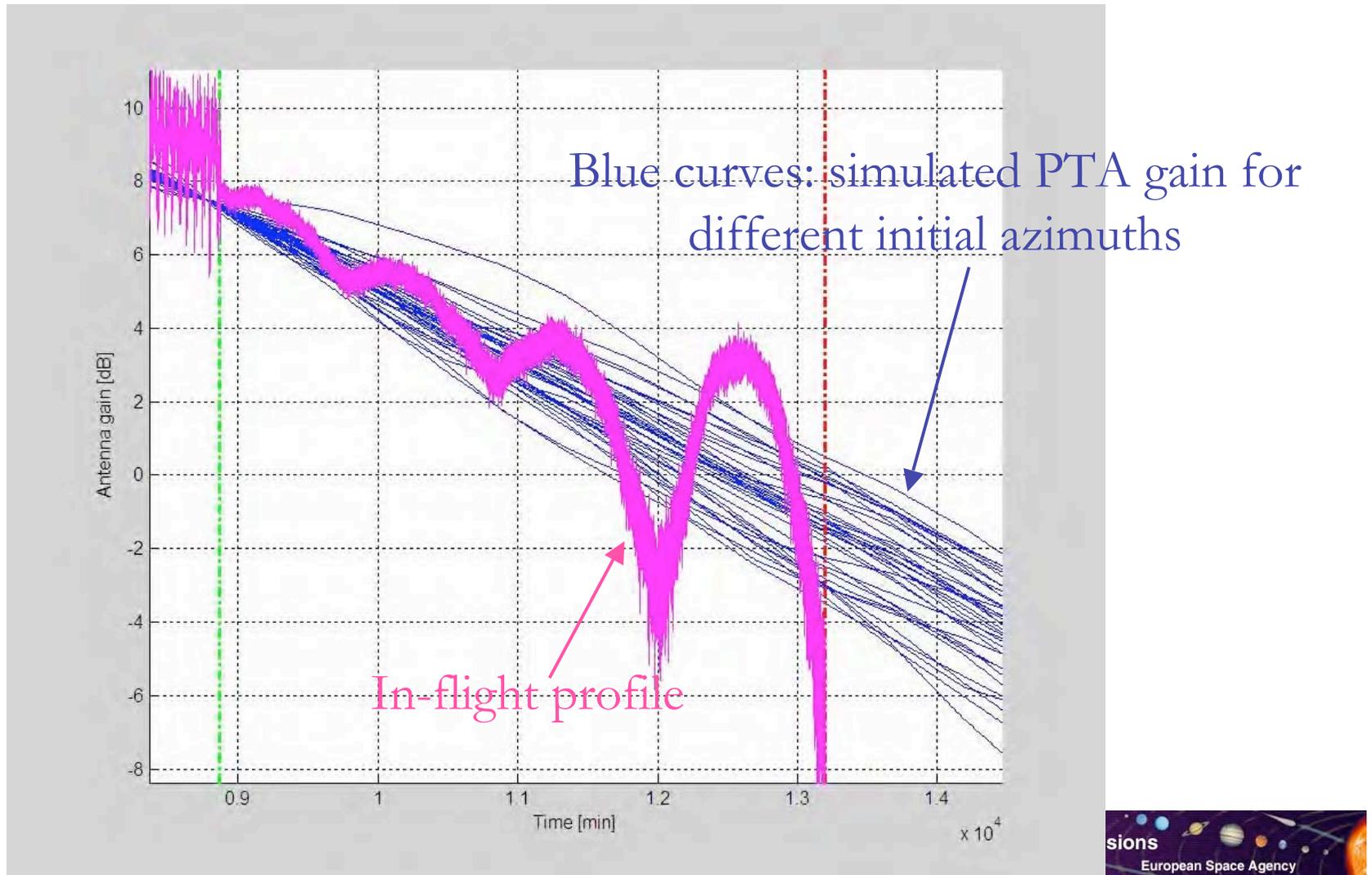
# RAY BENDING effect

**BENDING ANGLE - DEFOCUS LOSS  
FOR TITAN ATM-A PROFILE**

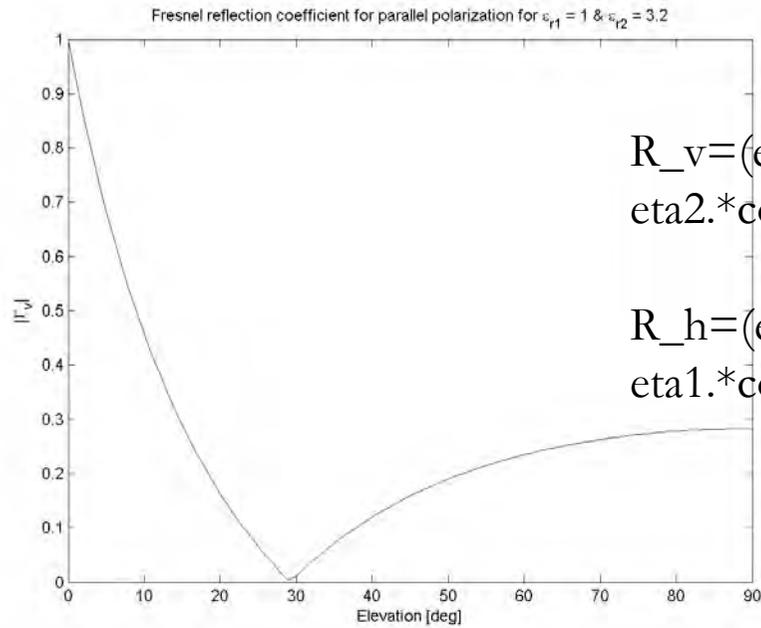


Oscillation not explained by 'SCANNING of the PTA pattern' (due to azimuth-elevation variation of orbiter position):

- too fast in time
- too large in power variation

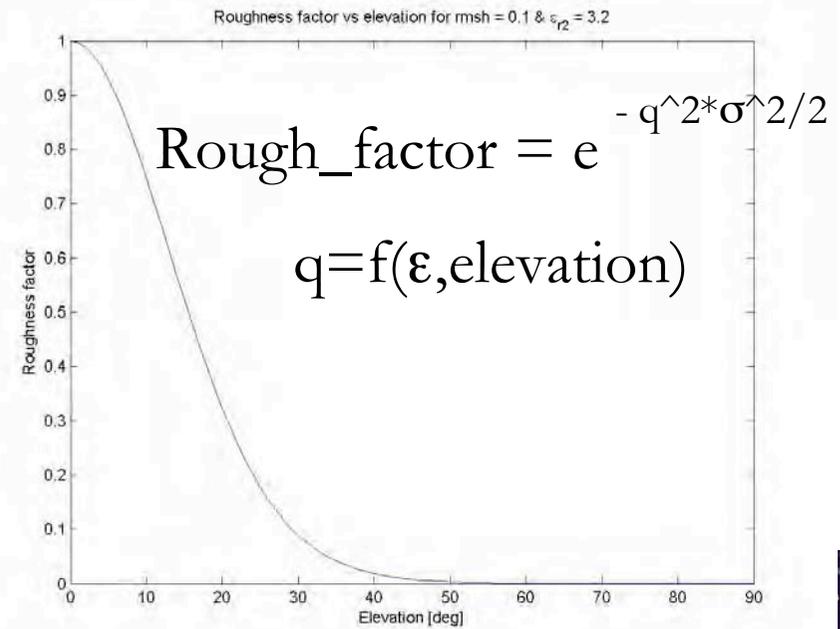
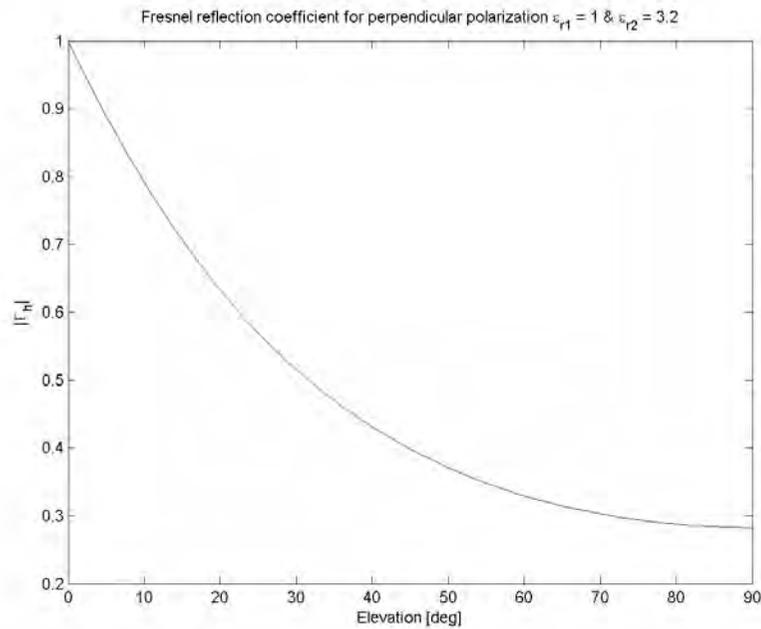


# Coefficients

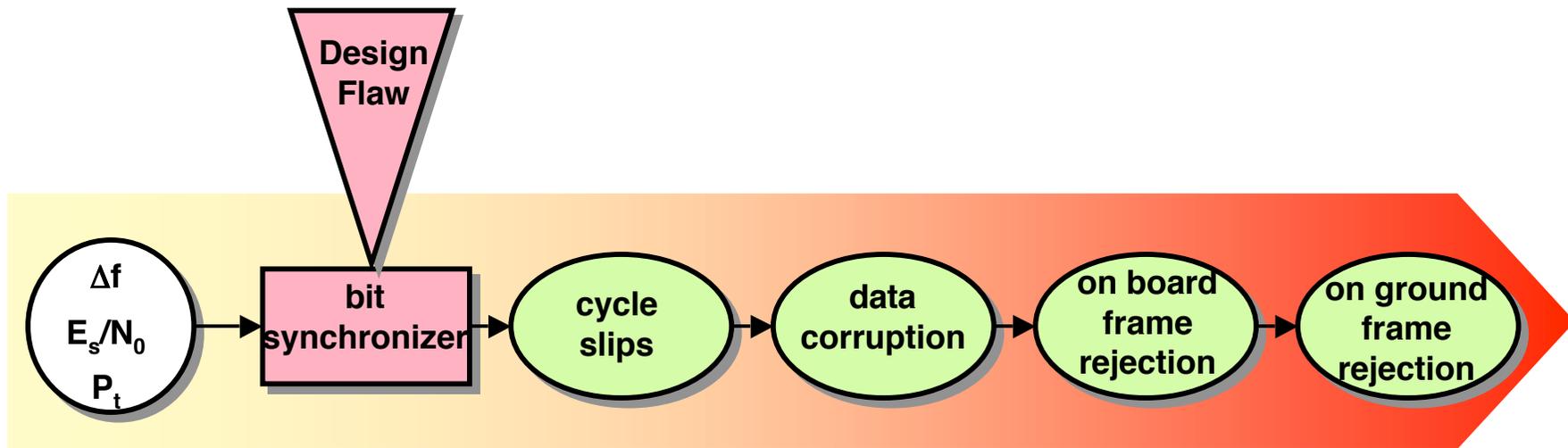


$$R_v = (\eta_1 \cdot \cos(\theta_1) - \eta_2 \cdot \cos(\theta_2)) / (\eta_1 \cdot \cos(\theta_1) + \eta_2 \cdot \cos(\theta_2));$$

$$R_h = (\eta_2 \cdot \cos(\theta_1) - \eta_1 \cdot \cos(\theta_2)) / (\eta_2 \cdot \cos(\theta_1) + \eta_1 \cdot \cos(\theta_2));$$



# Frame loss mechanism



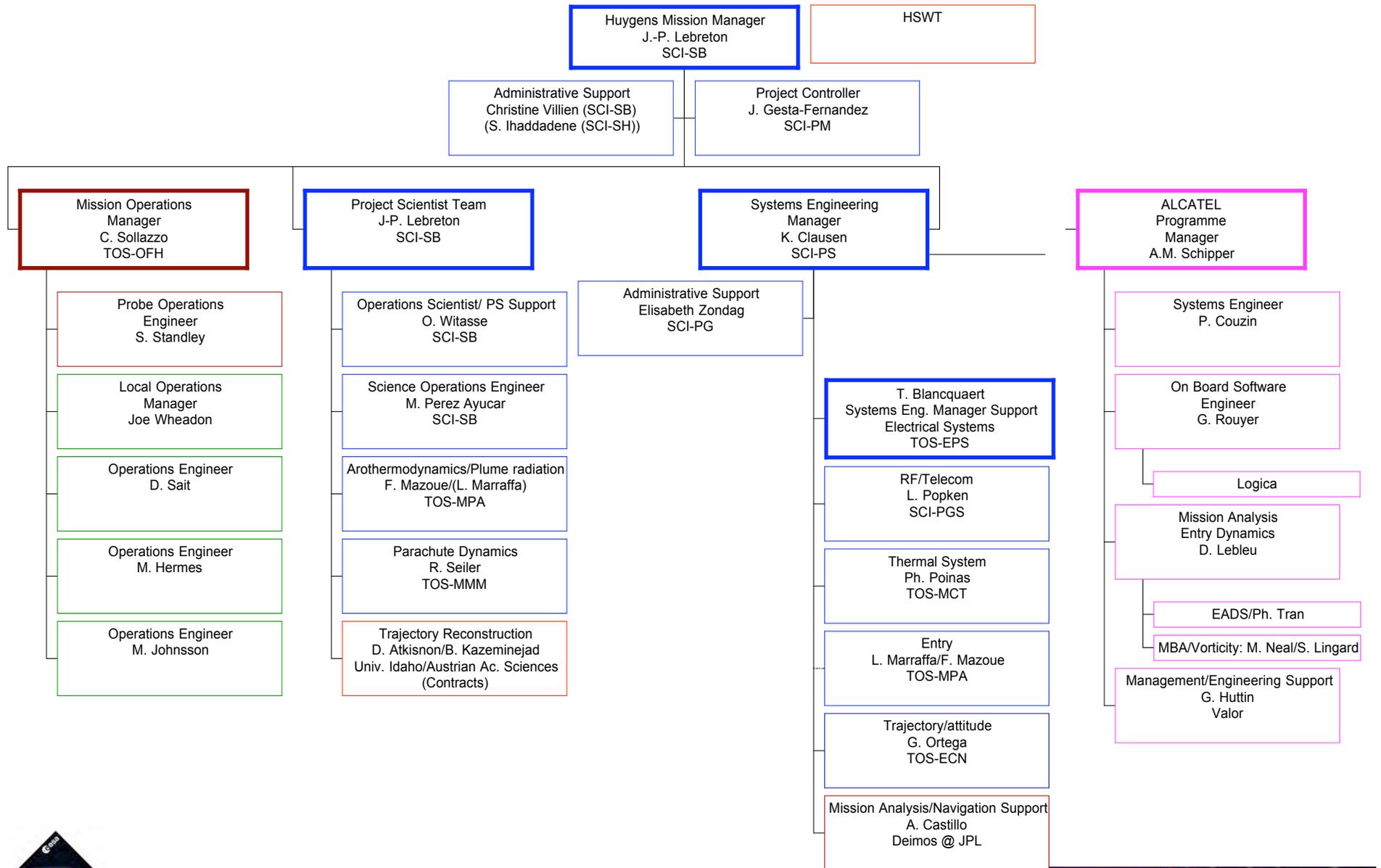
Relevant  
Parameters

The  
Problem

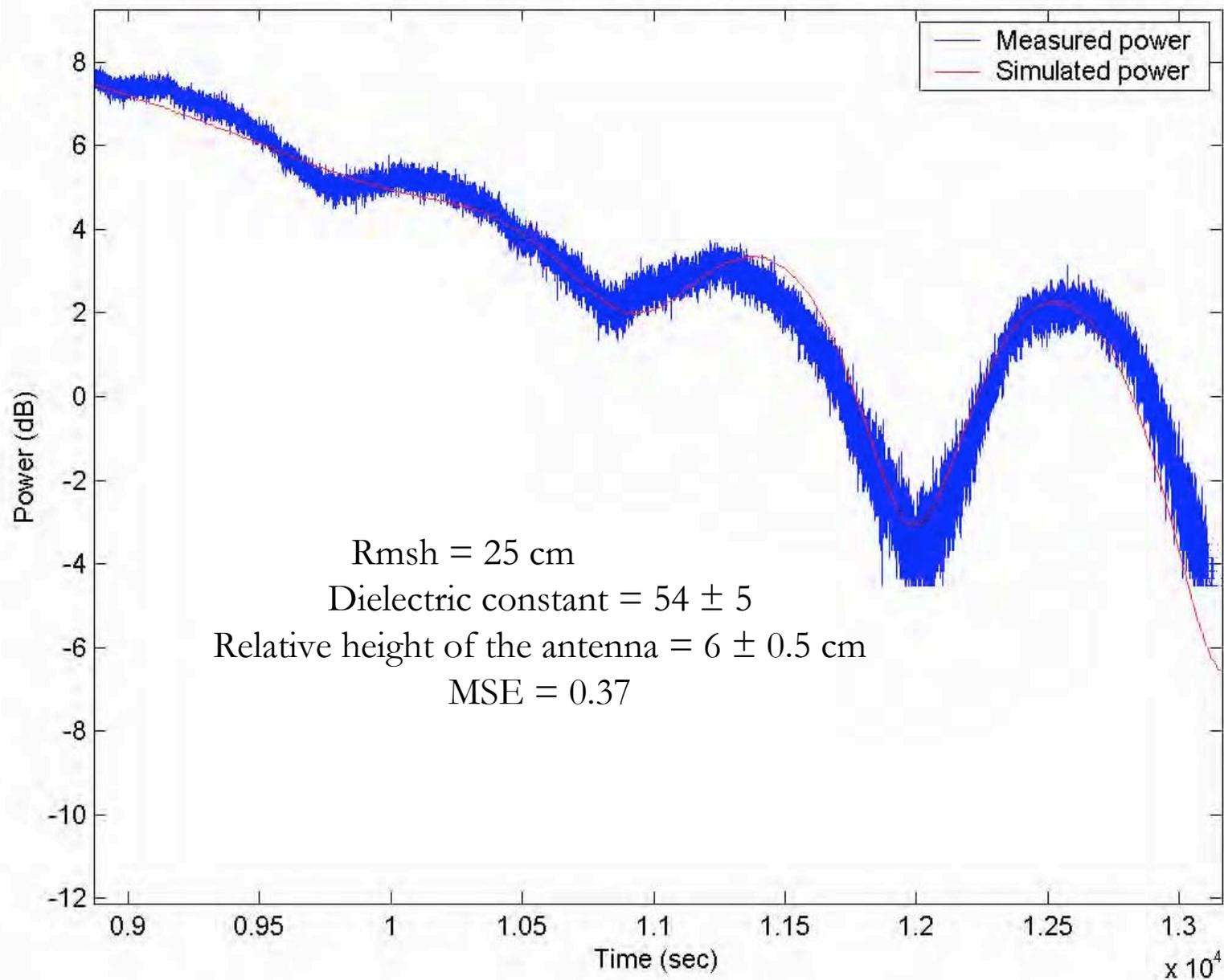
The  
Effect



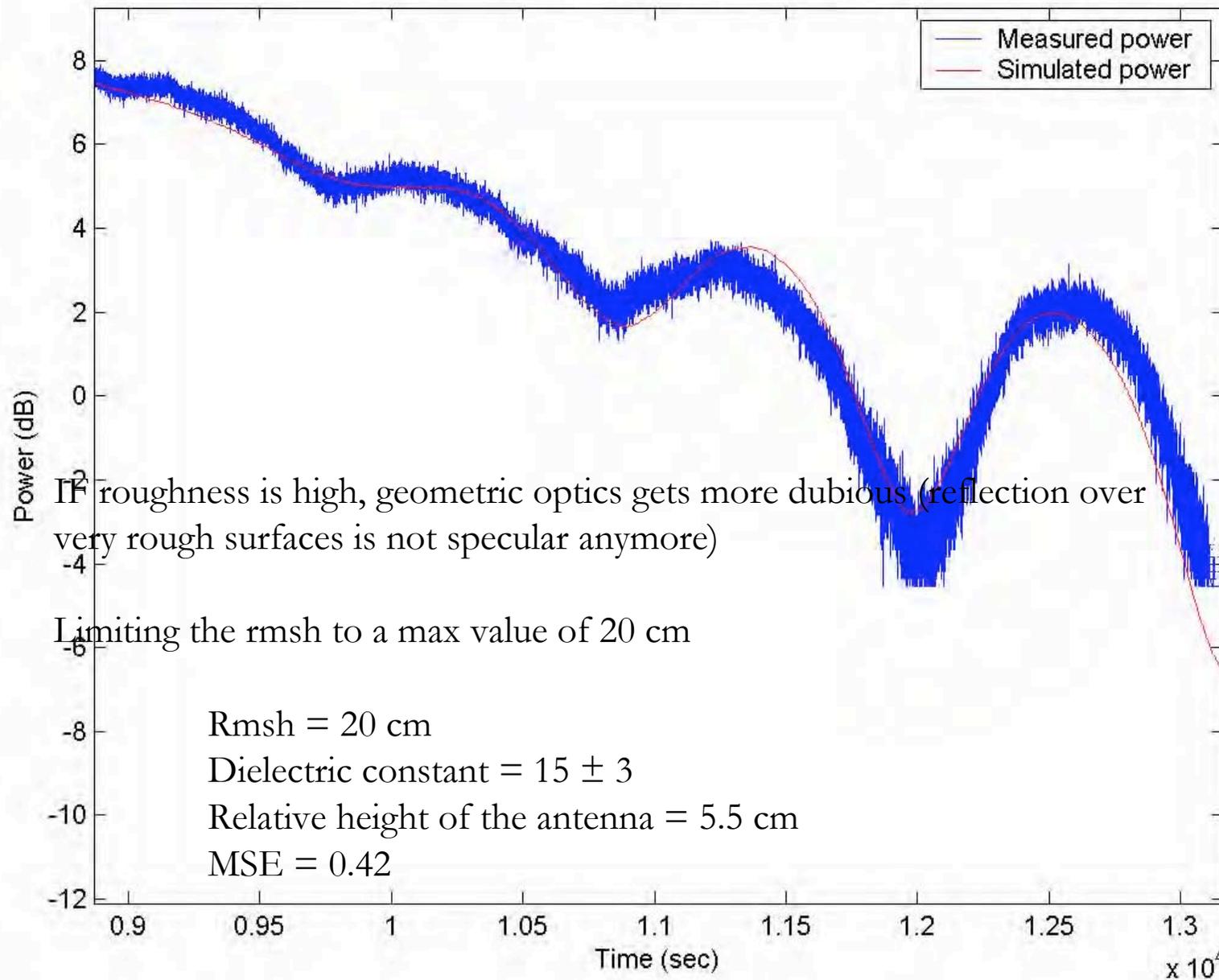
# ESA Huygens Mission Team: HMT



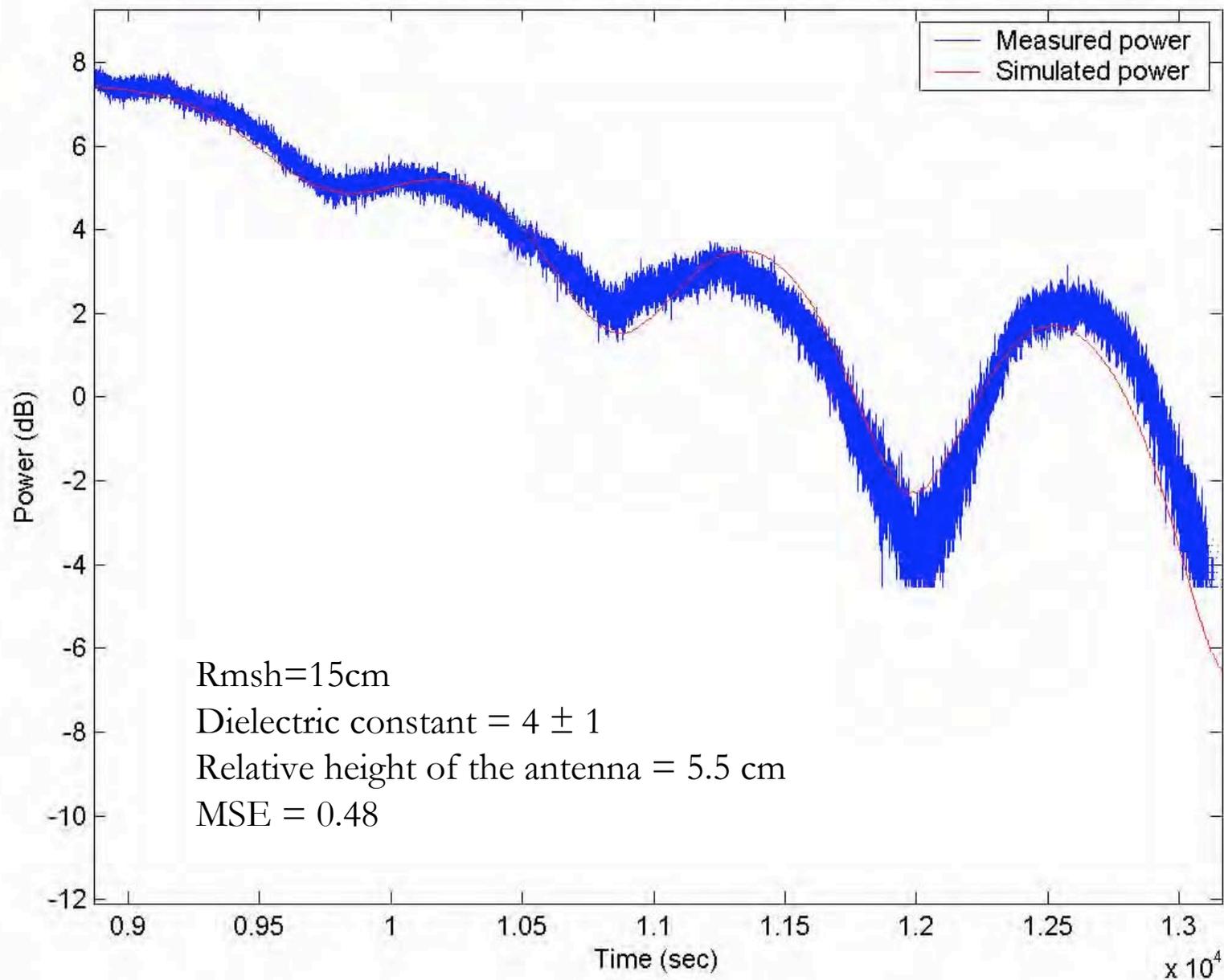
Measured power vs time. Simulated power vs time for  $\epsilon_r = 54$ ,  $r_{msh} = 0.25$ ,  $h_r = 0.06$



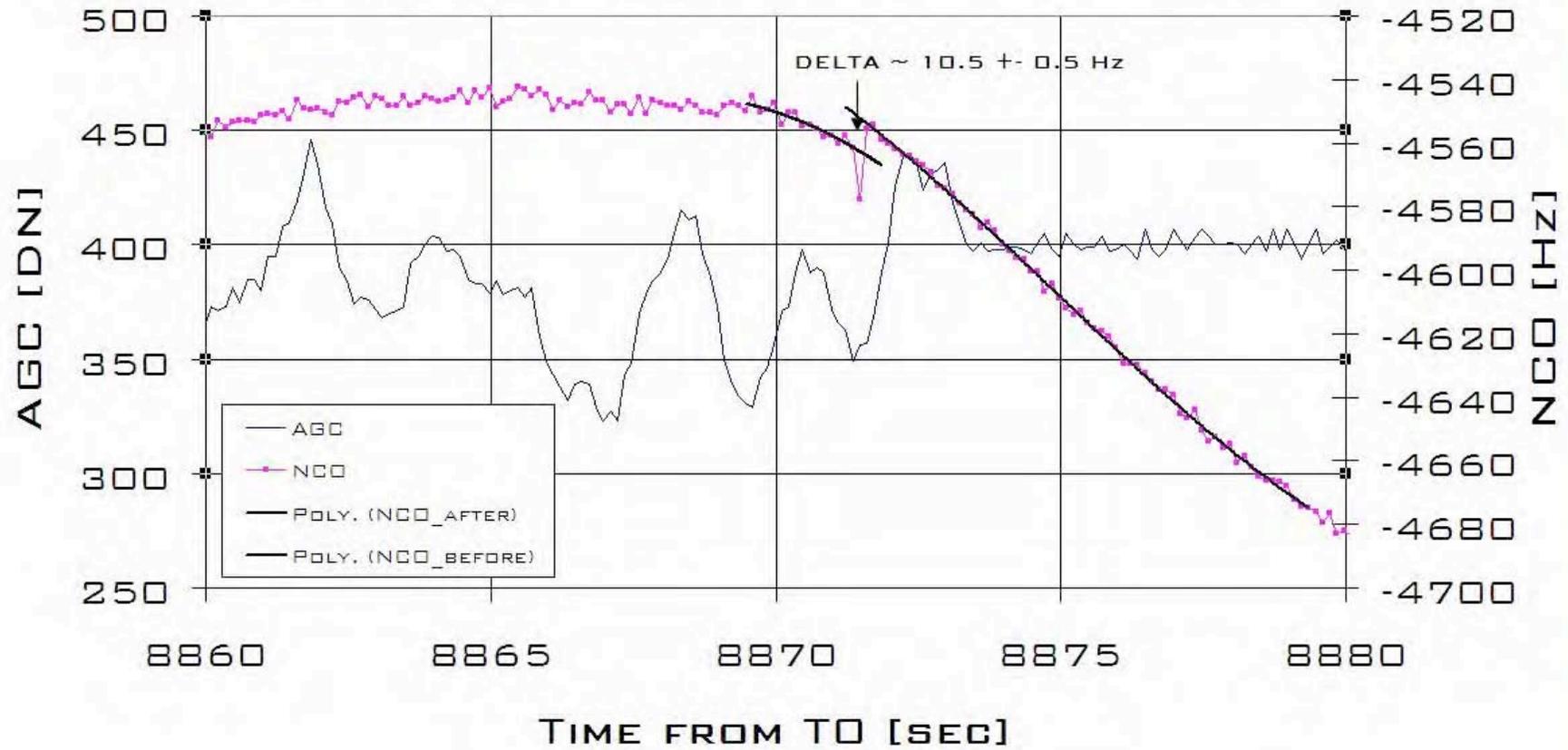
Measured power vs time. Simulated power vs time for  $\epsilon_r = 15$ ,  $\text{rms}_h = 0.2$ ,  $h_r = 0.055$



Measured power vs time. Simulated power vs time for  $\epsilon_r = 4$ ,  $r_{msh} = 0.15$ ,  $h_r = 0.055$

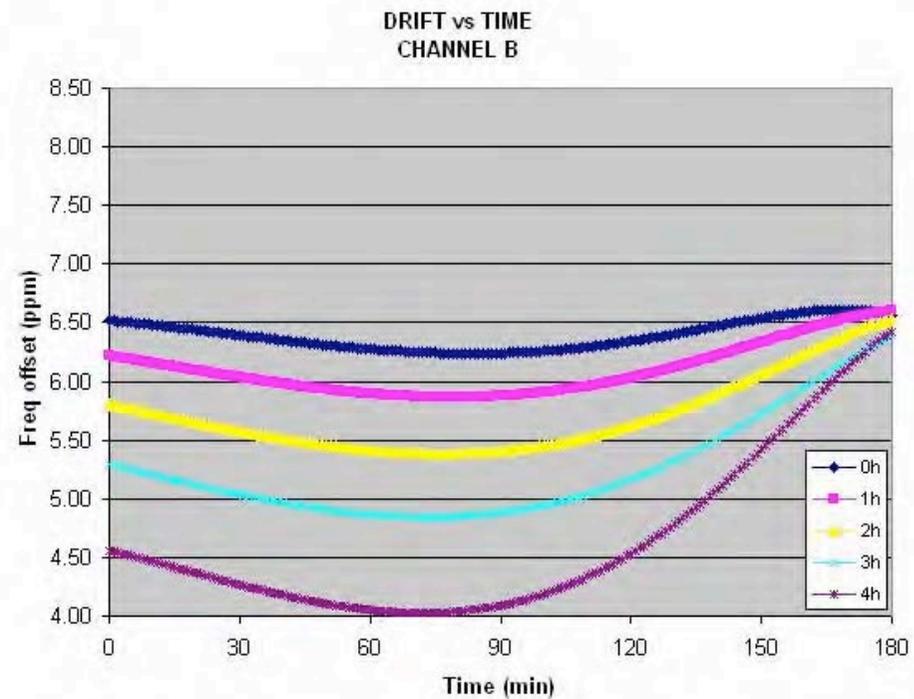
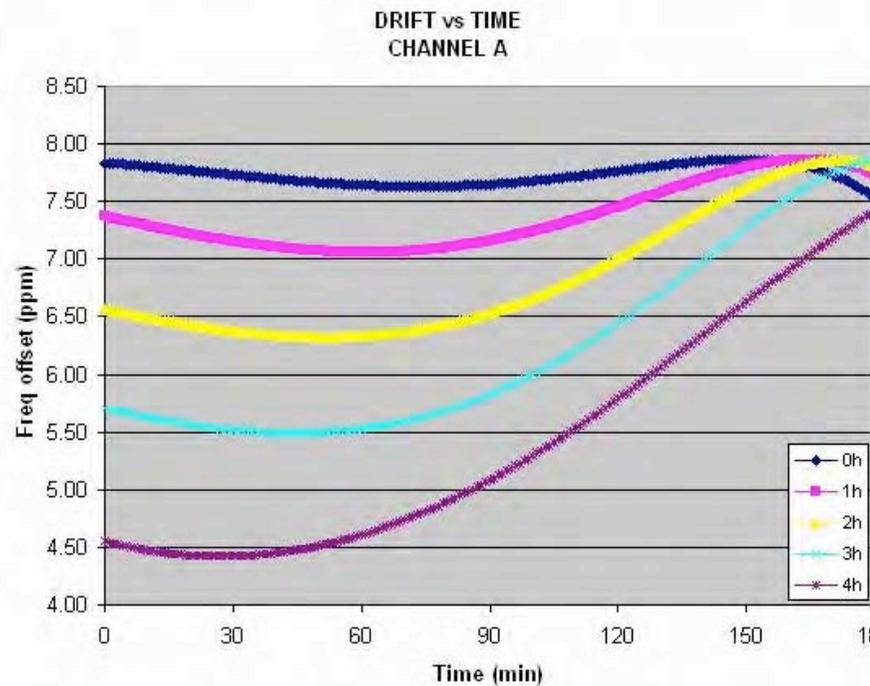


# HUYGENS RADIOLINK FREQUENCY AND POWER AROUND IMPACT

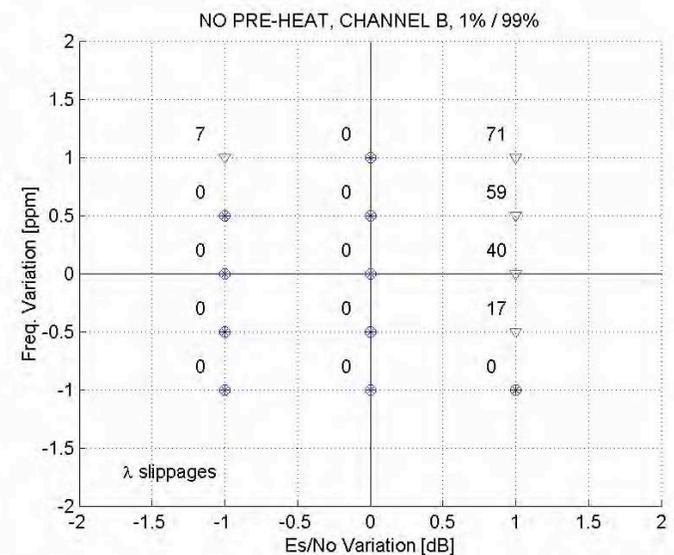
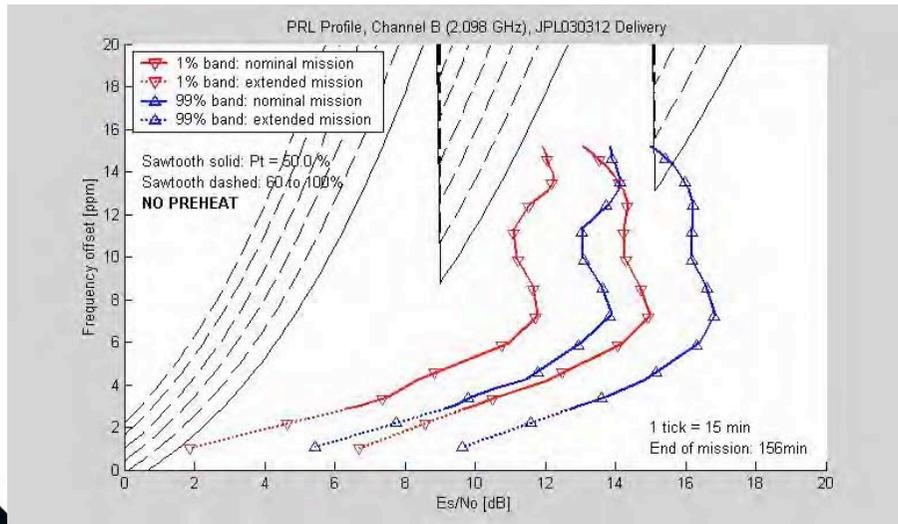
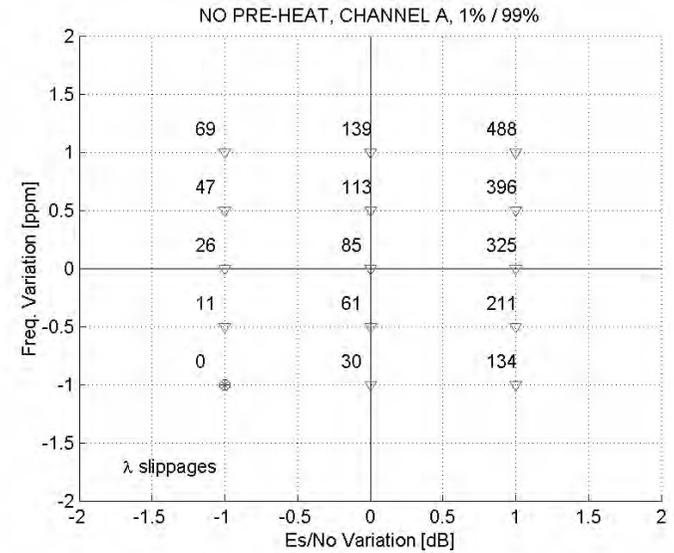
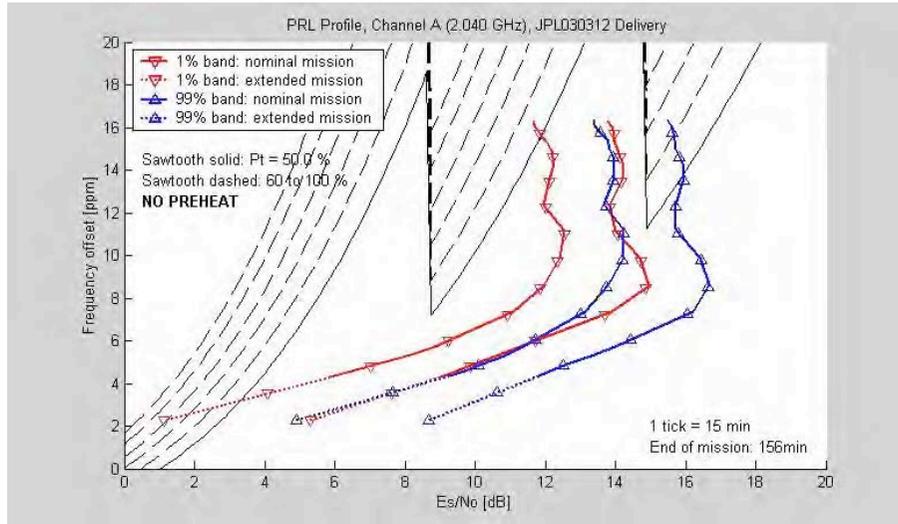


## PRE-HEATING 4 hours:

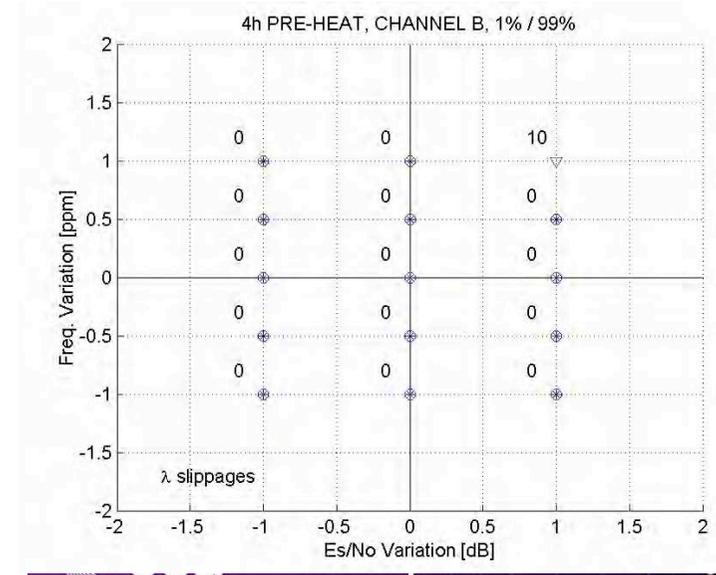
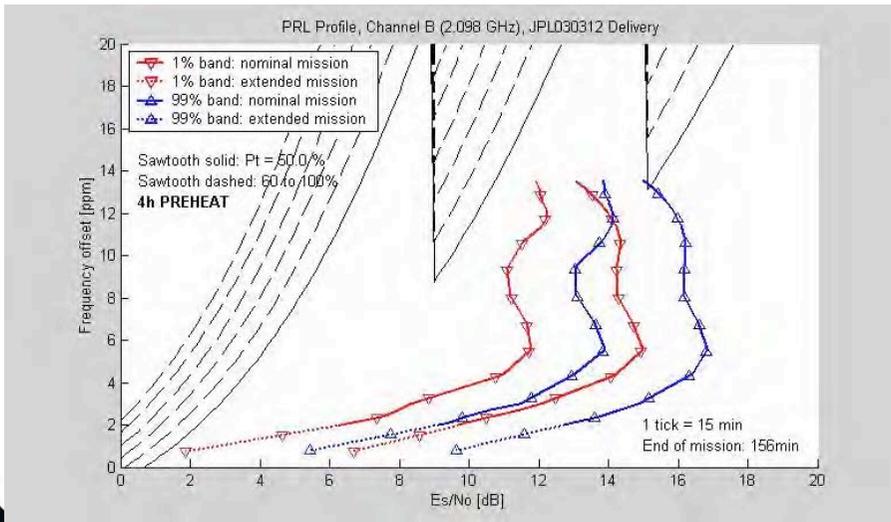
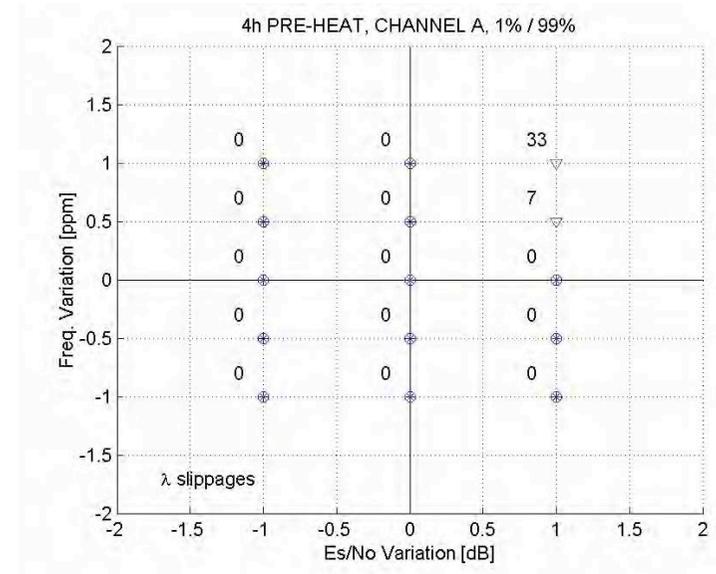
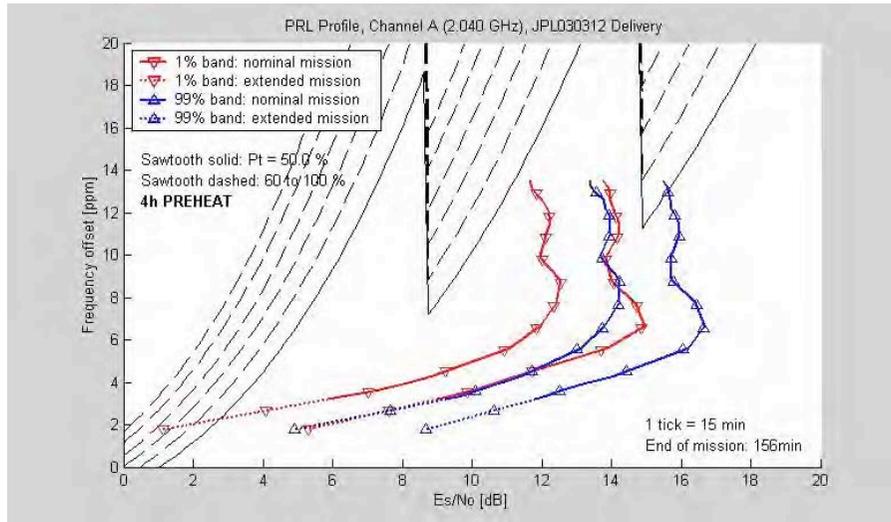
Warm-up the quartz oscillator generating the frequency for the transmitted data frequency, to shift it and compensate the Doppler



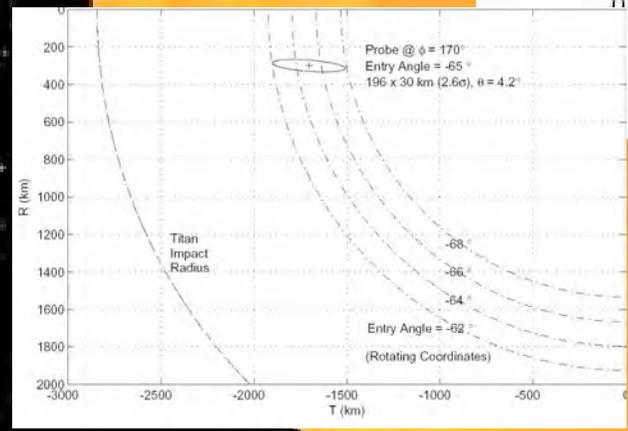
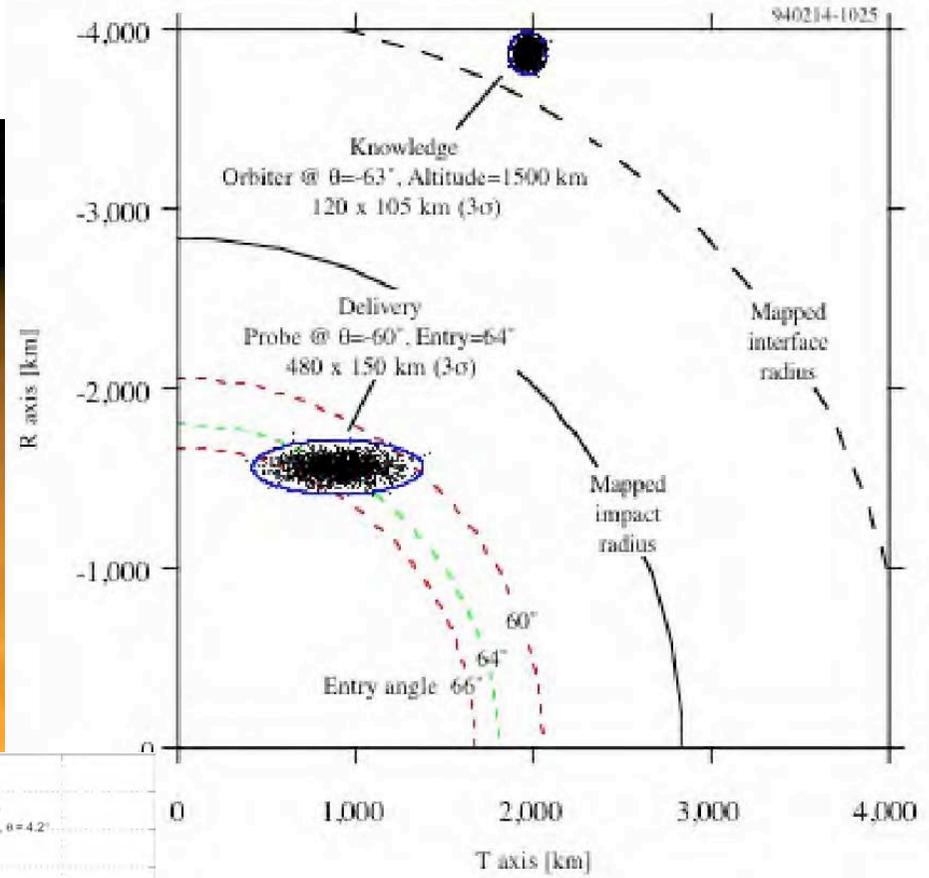
# Robustness: NO PRE-HEAT



# Robustness: 4h PRE-HEAT (implemented option)



# TARGETING



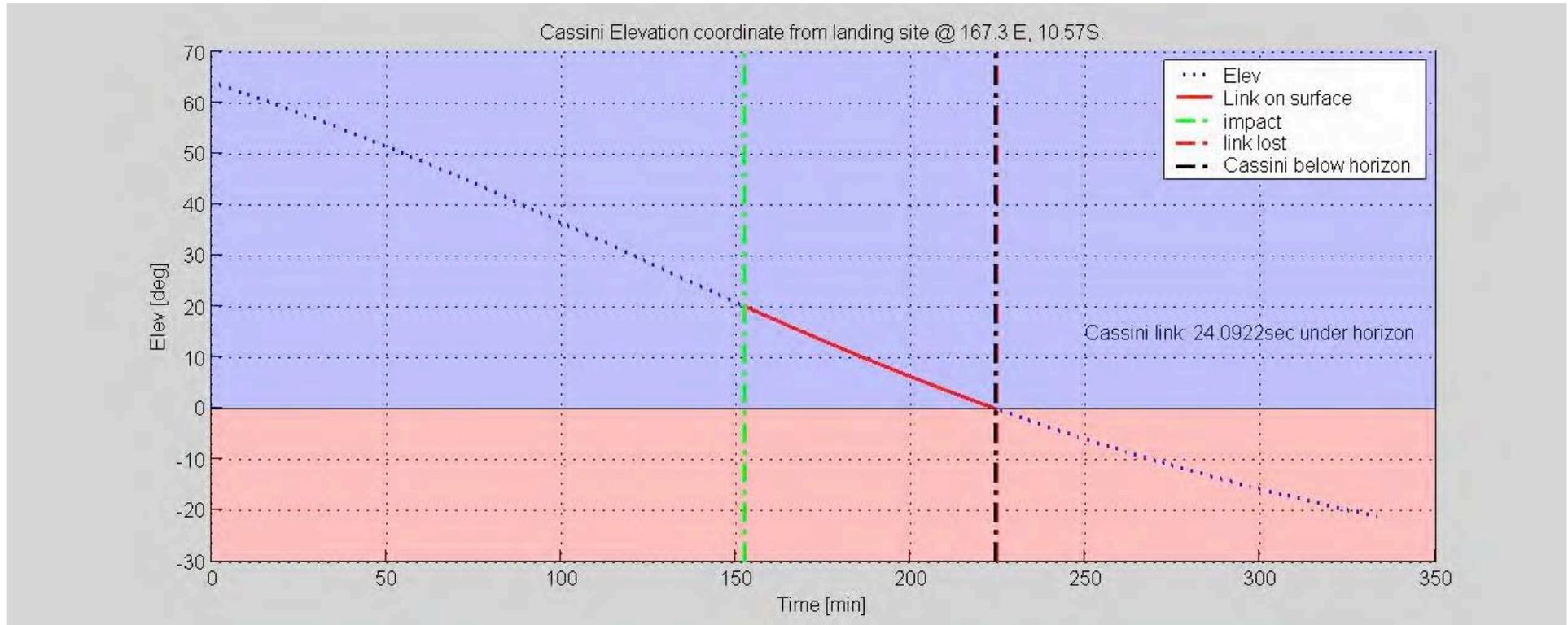
NUEVA GEOMETRIA

VIEJA GEOMETRIA



## END OF LINK: Cassini ELEV vs TIME

- Cassini very close to disappear under the horizon at the time of loss of link (**24 seconds**) !!

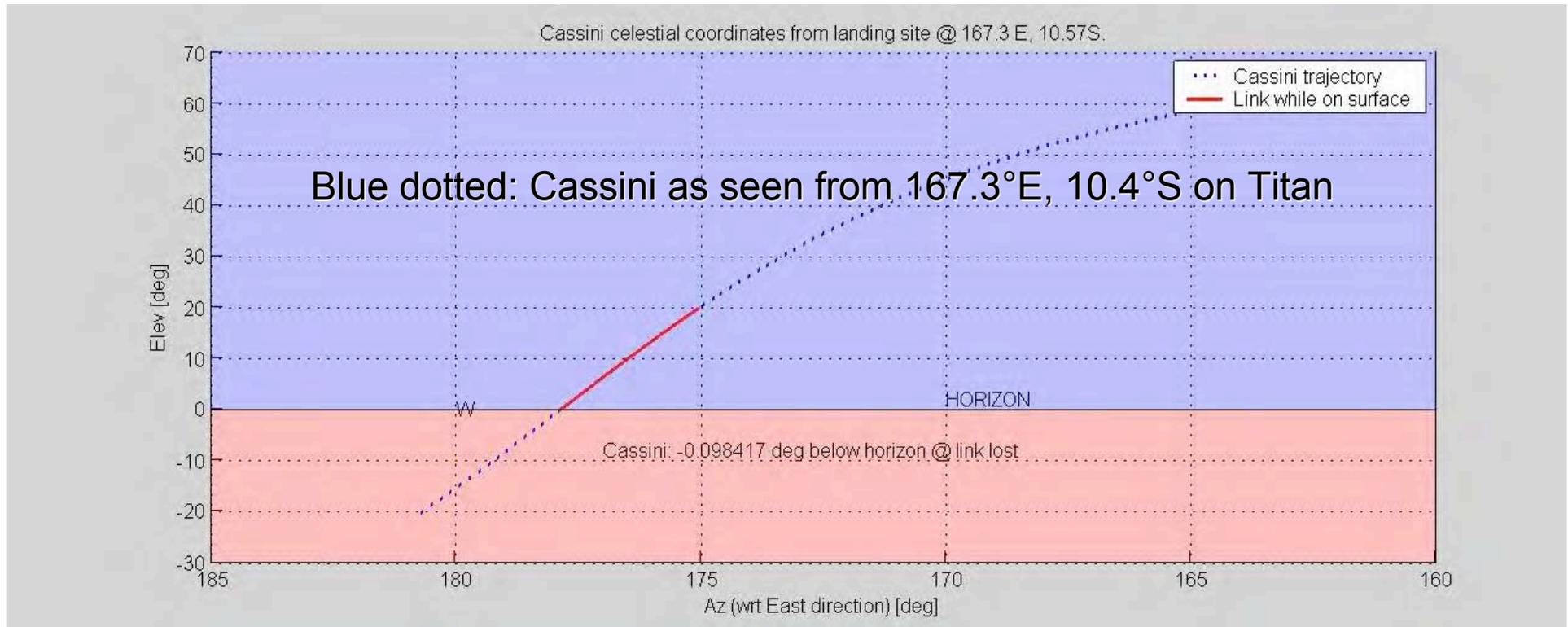


Blue dotted: Cassini ELEV as seen from  $167.3^{\circ}$   $-10.4^{\circ}$  on Titan  
RED: Link ON SURFACE



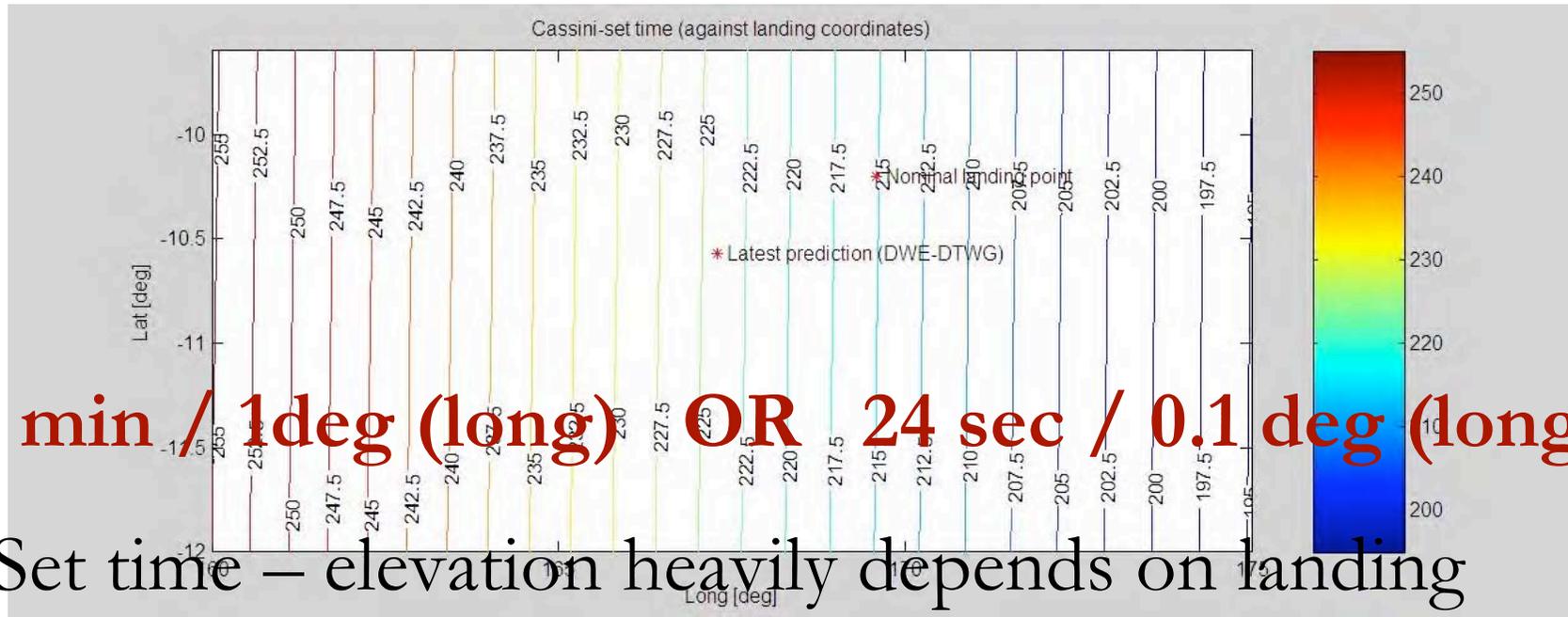
# END OF LINK: CASSINI AZ-ELEV FROM LANDING SITE

Heavily depending on landing coordinates  
Waiting for update of HUYGENS landing site



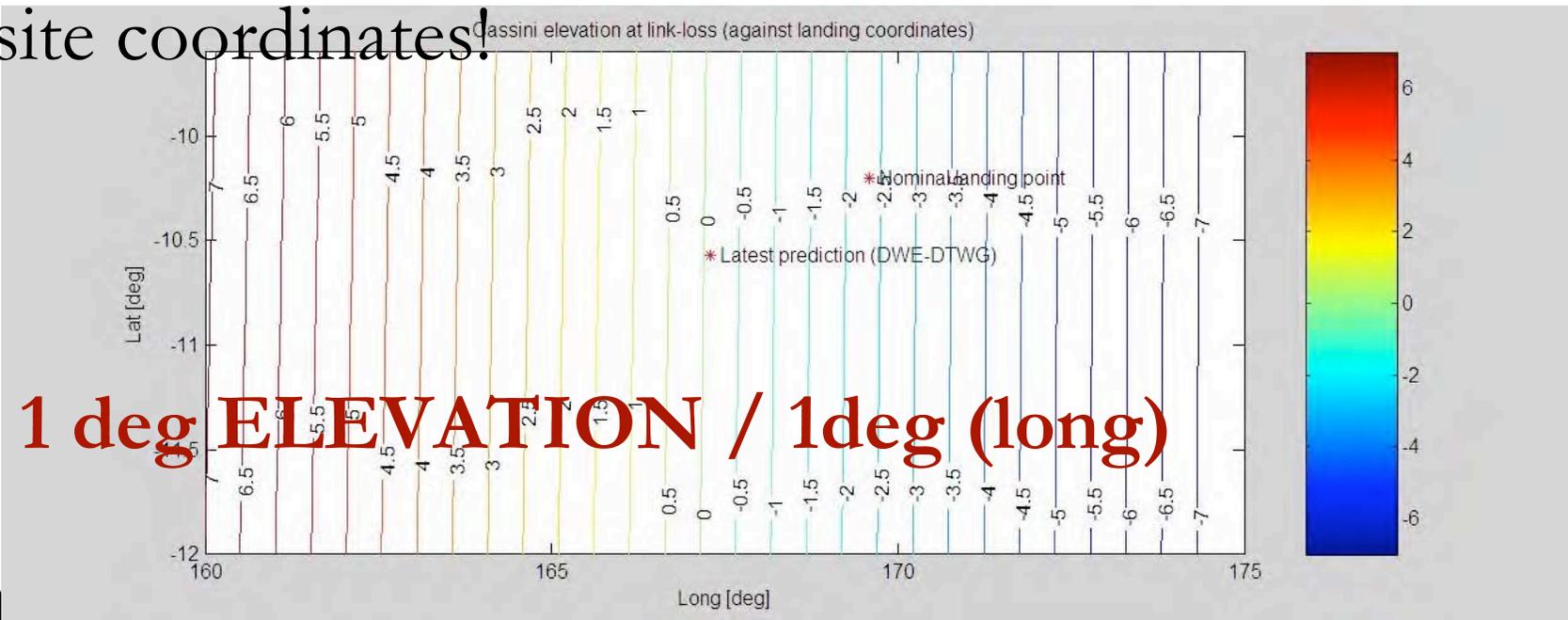
Grazing bending of ray ~ 1deg max  
Defocussing loss ~ 10%





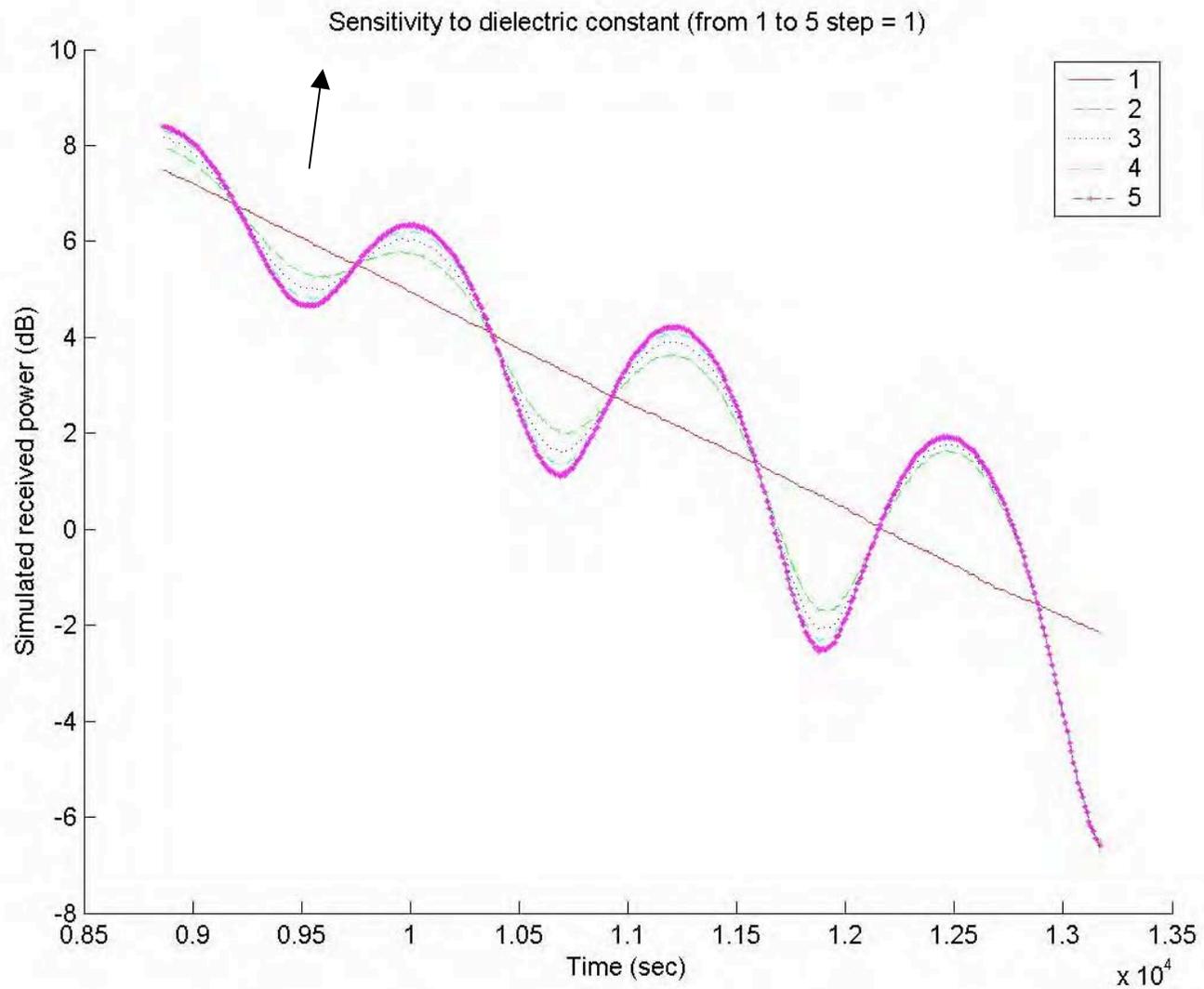
**4 min / 1deg (long) OR 24 sec / 0.1 deg (long)**

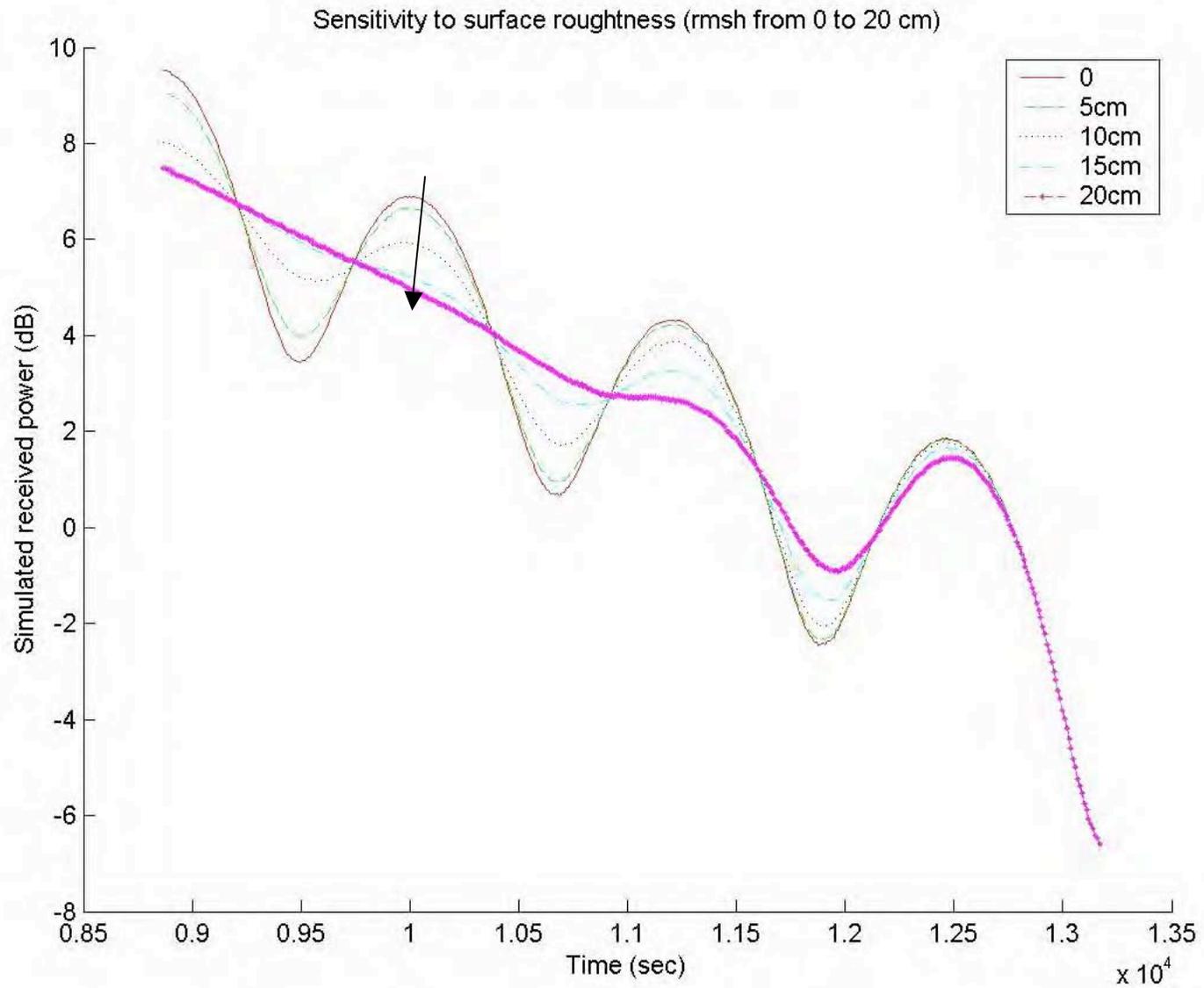
Set time – elevation heavily depends on landing site coordinates!



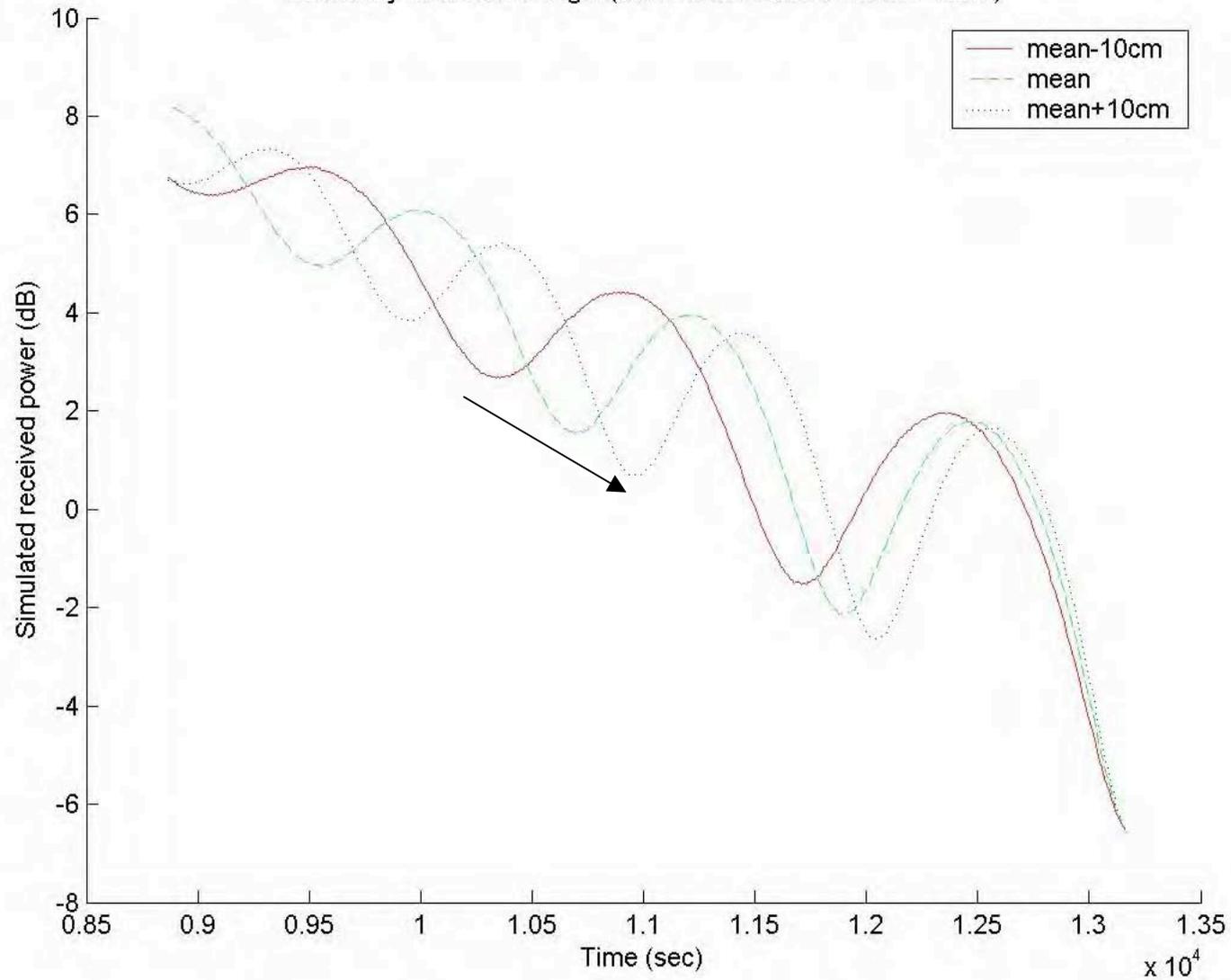
**1 deg ELEVATION / 1deg (long)**



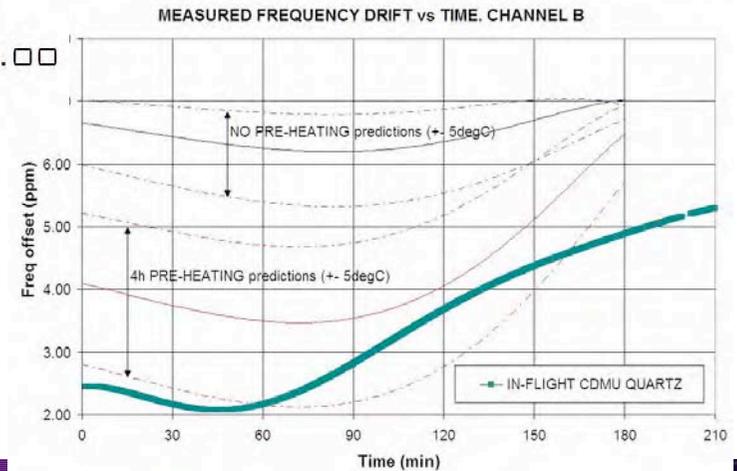
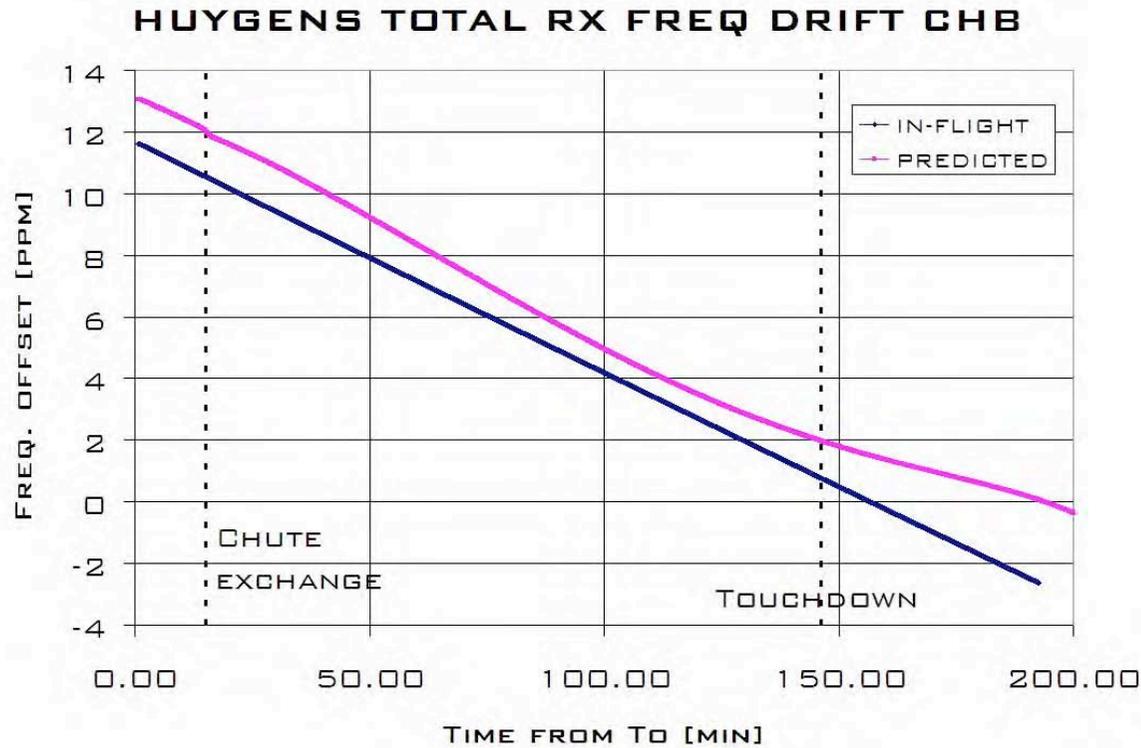




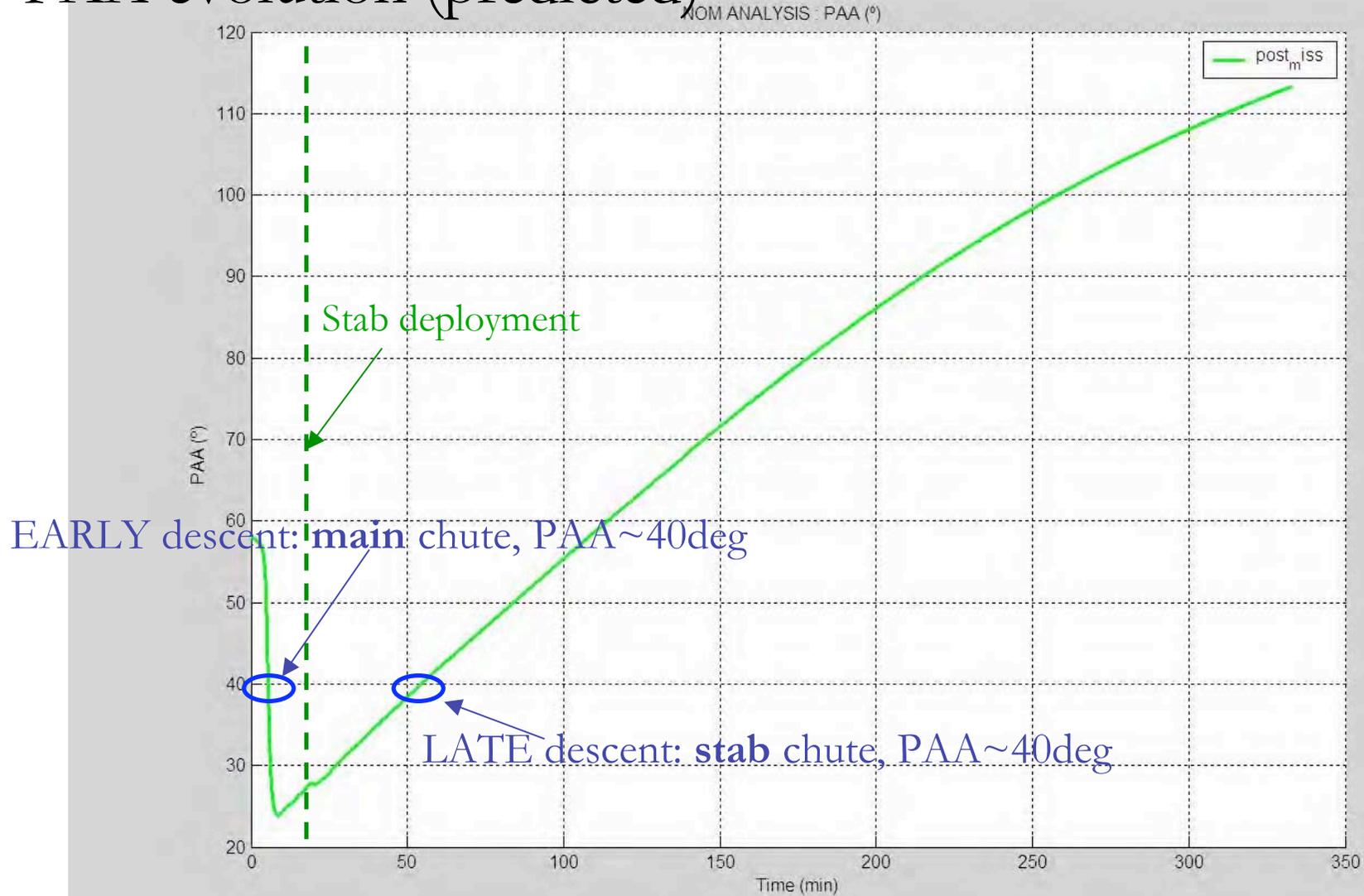
Sensitivity to antenna height (from mean-10cm to mean+10cm)



# The reconstructed received frequency profile

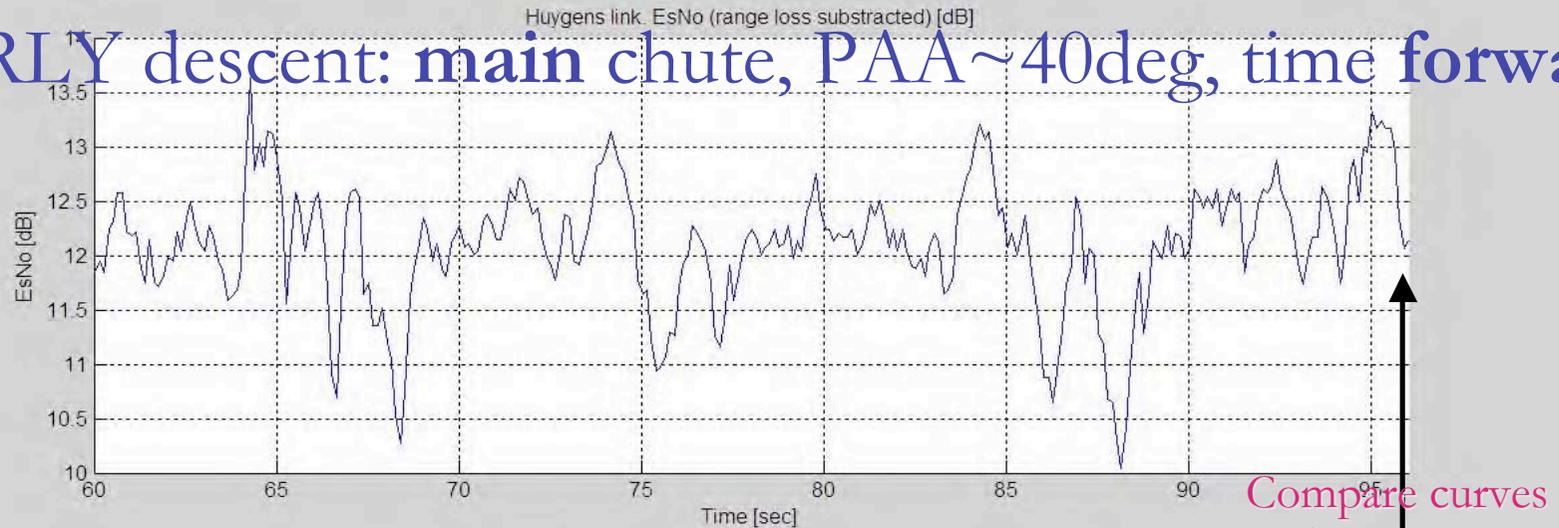


# PAA evolution (predicted)

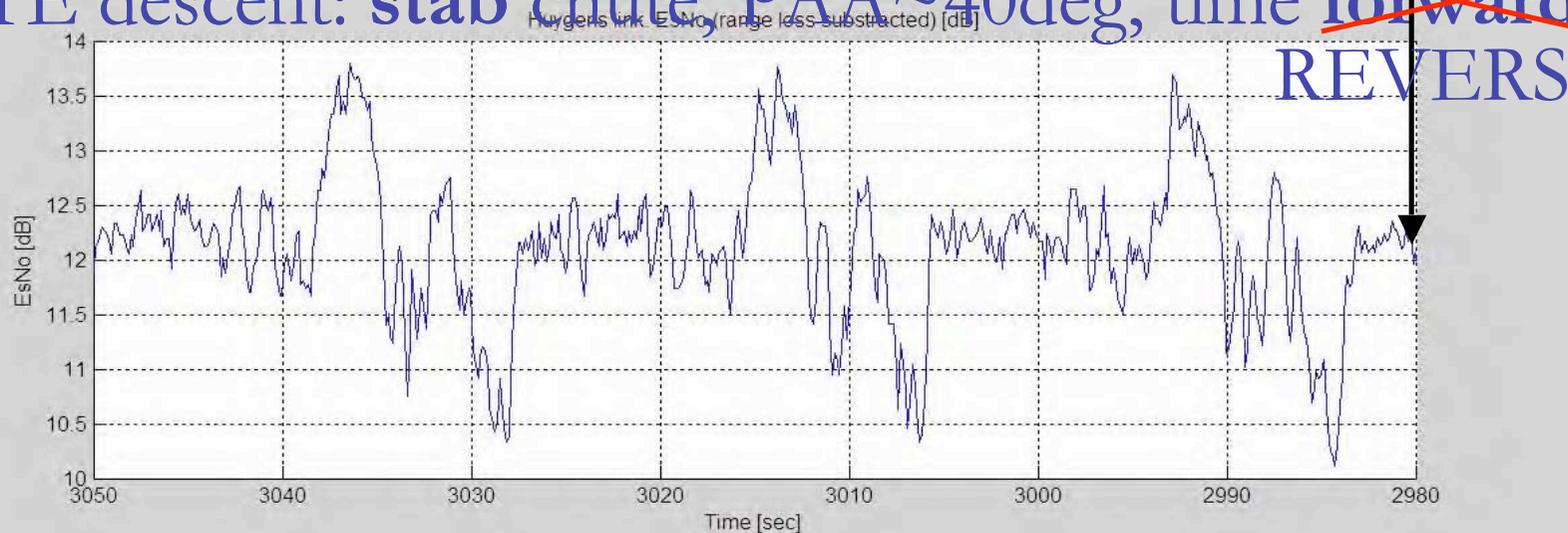


# SPIN REVERSAL – MATCHING EARLY DESCENT with LATE DESCENT

EARLY descent: main chute, PAA~40deg, time forward



LATE descent: stab chute, PAA~40deg, time ~~forward~~  
REVERSED



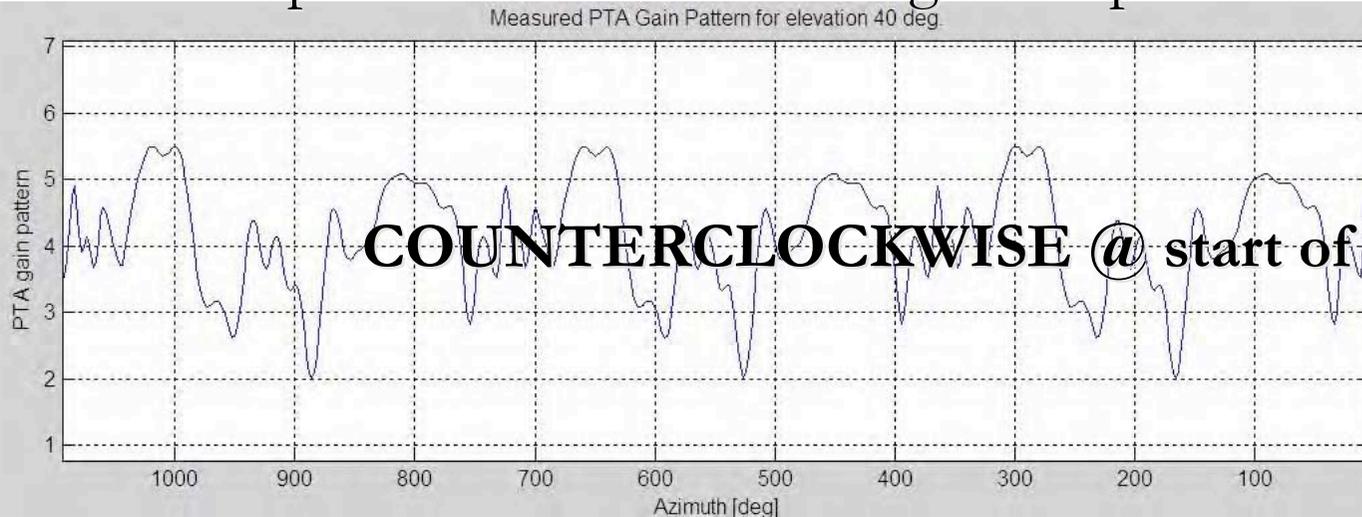
# SPIN REVERSAL – MATCHING THE TESTED PTA PATTERNS

Link mission

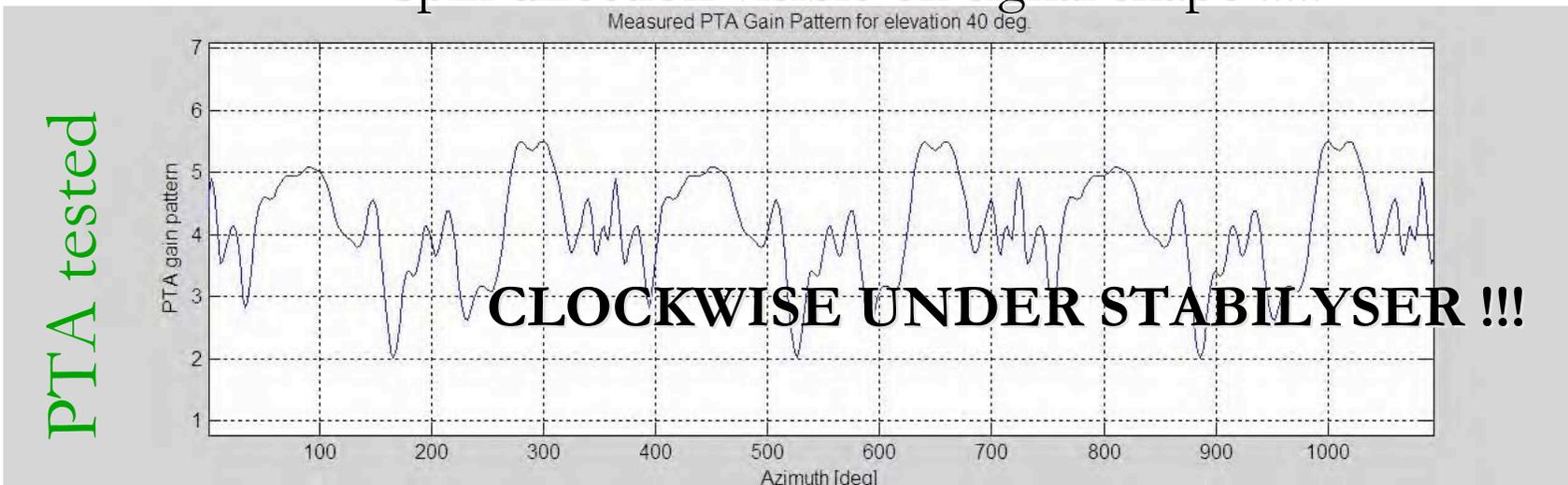
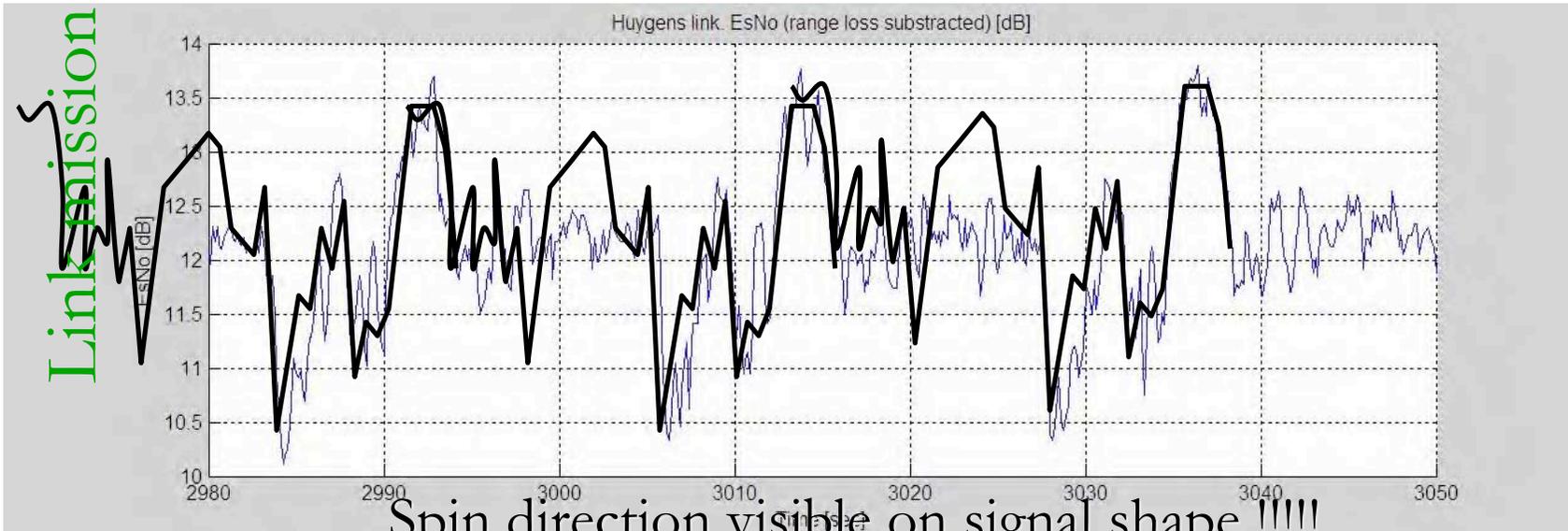


Spin direction visible on signal shape !!!!!

PTA tested



# SPIN REVERSAL – MATCHING THE TESTED PTA PATTERNS



Oscillation not explained by 'SCANNING of the PTA pattern' (due to azimuth-elevation variation of orbiter position):

- too fast in time
- too large in power variation

